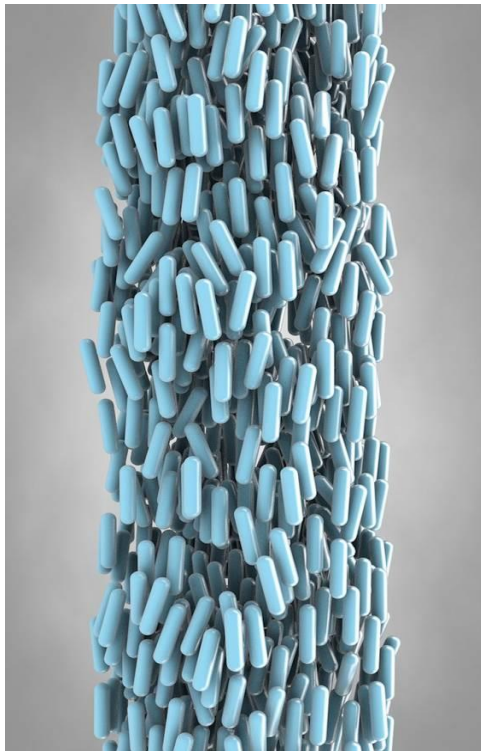
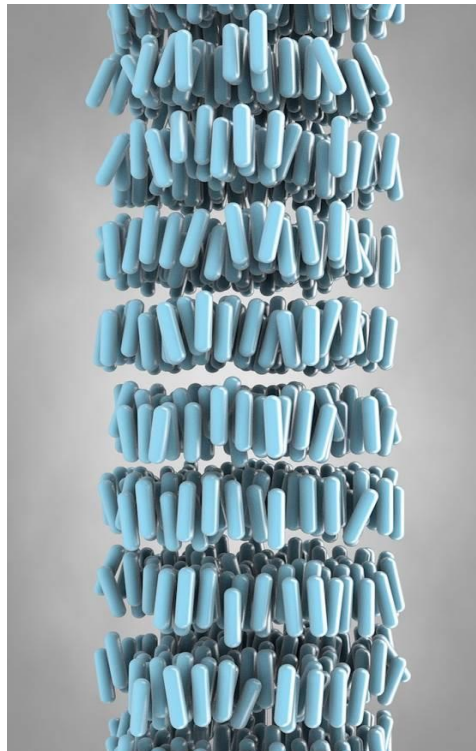


Overview of lecture 1

- The self-organization of rod-like molecules gives rise to hybrid states or mesophases, that are half-liquid (they can flow) and half-solid (they have anisotropic properties) = liquid crystals.



Nematic



Smectic A



Cholesteric

Overview of lecture 1

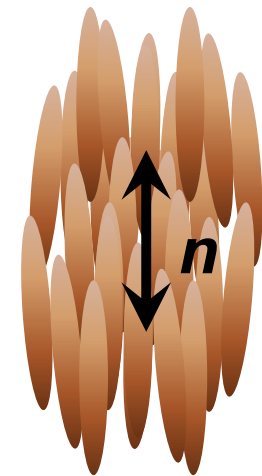
- The simplest mesophase is the nematic phase, in which there is no positional order and a long-range orientational order: the average local orientation of the molecular axis is given by the director field \mathbf{n} .
- The degree of orientational order is described by the scalar order parameter S

$$S = \langle P_2(\cos \theta) \rangle = \left\langle \frac{3}{2} \cos^2 \theta - \frac{1}{2} \right\rangle$$

- Anisotropy \Rightarrow Physical properties are different along \mathbf{n} and orthogonally to it.

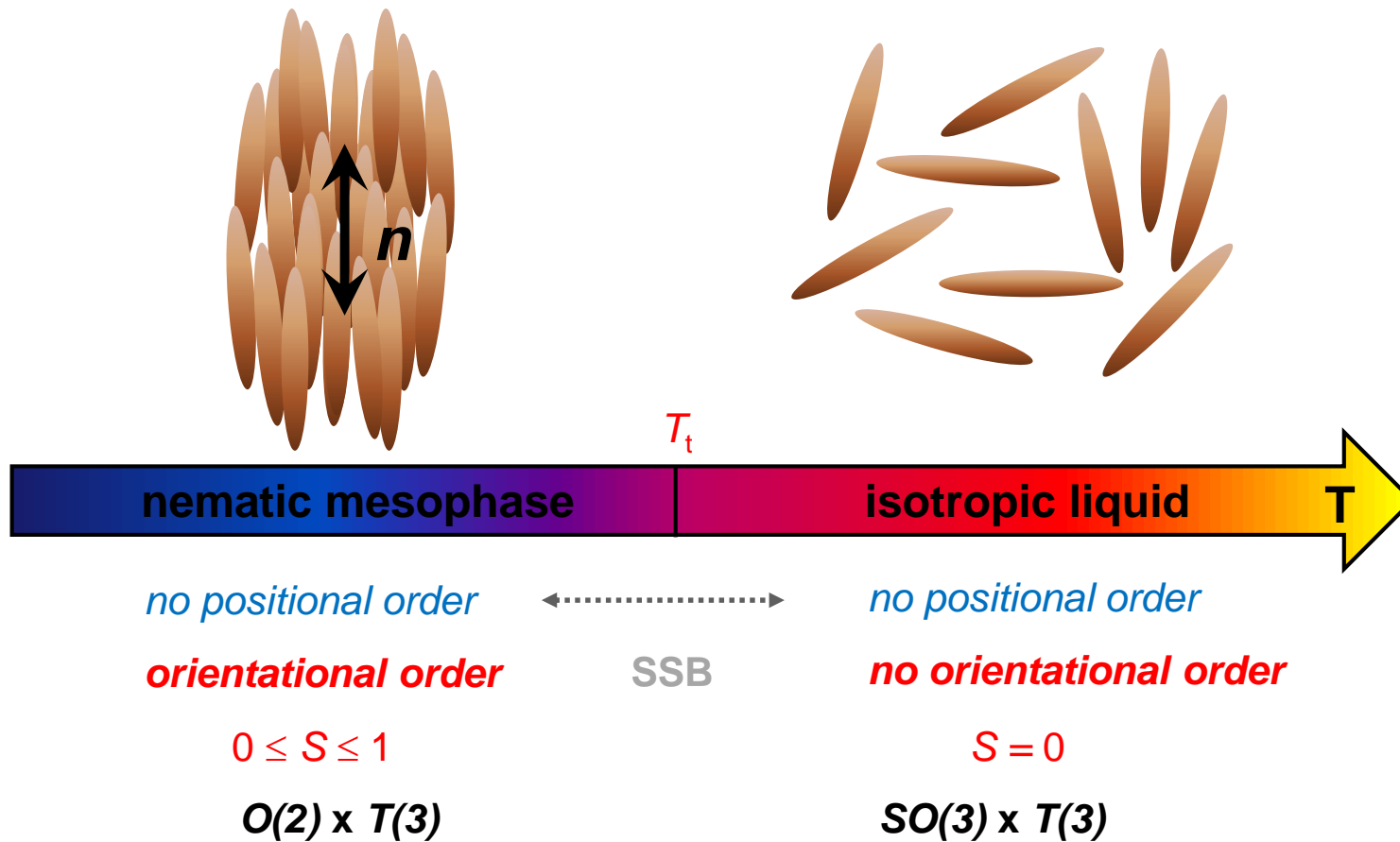
Example: optics

$$\mathbf{D} = \overset{=}{\boldsymbol{\varepsilon}} \mathbf{E} \quad \overset{=}{\boldsymbol{\varepsilon}} = \begin{pmatrix} \varepsilon_{\perp} & 0 & 0 \\ 0 & \varepsilon_{\perp} & 0 \\ 0 & 0 & \varepsilon_{\parallel} \end{pmatrix}$$



Overview of lecture 1

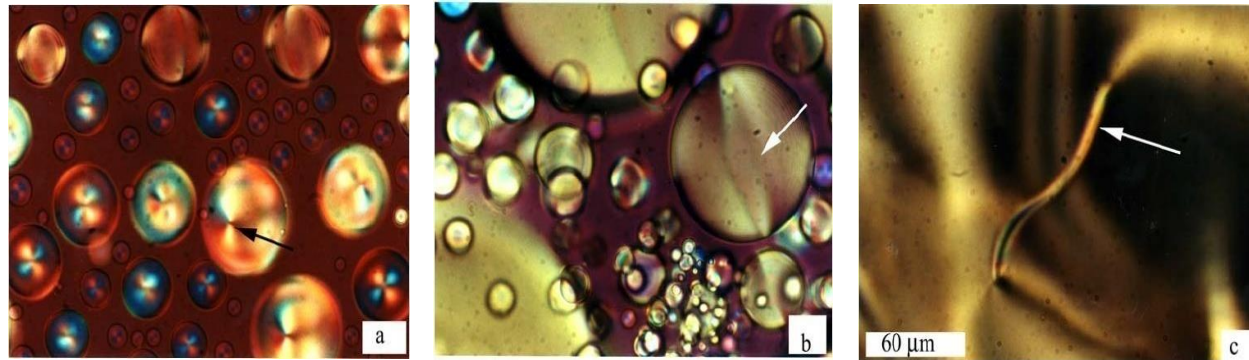
- The isotropic-nematic phase transition presents a spontaneous symmetry breaking:



Overview of lecture 1

M Kleman et al. Phil Mag 86, 4117 (2006)

- The isotropic-nematic phase transition presents a spontaneous symmetry breaking leading to defect production:



isotropic liquid

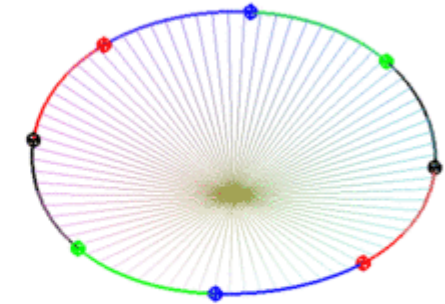
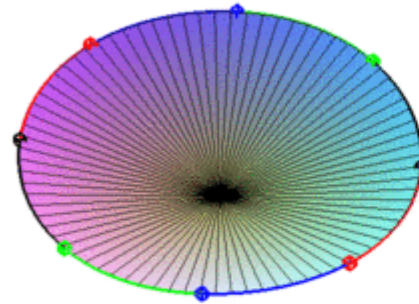
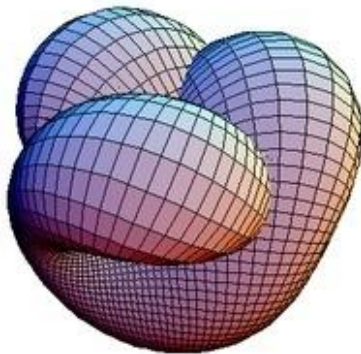
Domino cascade = SSB



nematic mesophase
with a defect

Overview of lecture 1

- The order parameter space is the real projective plane $M = SO(3)/O(2) \sim S^2 / \mathbb{Z}_2 \leftrightarrow$ « Boy surface »



- The content of homotopy groups predicts the kind of defects that may appear in the mesophase:

$$\pi_0(M) = \mathbb{I}$$

No domain wall

$$\pi_1(M) = \mathbb{Z}_2$$

Line defects: disclinations...

$$\pi_2(M) = \mathbb{Z}$$

Monopoles: hedgehogs...

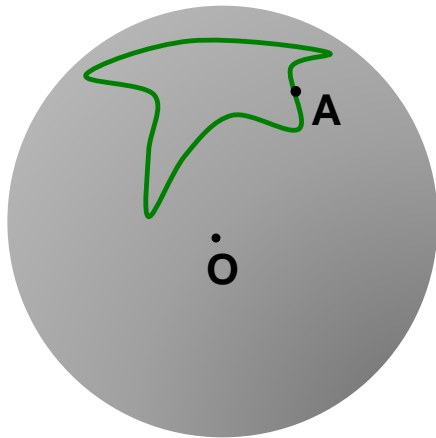
$$\pi_3(M) = \mathbb{Z}$$

Textures: skyrmions...

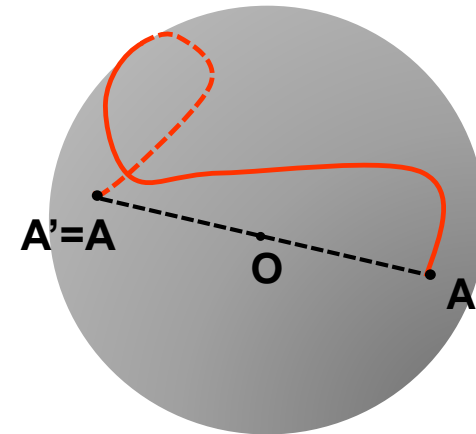
Overview of lecture 1

- The content of the first homotopy group means that there are only two equivalence classes for the linear defects:

unstable



$$\pi_1(\mathbb{R}P^2) = \mathbb{Z}_2 = \{0, 1\}$$



stable

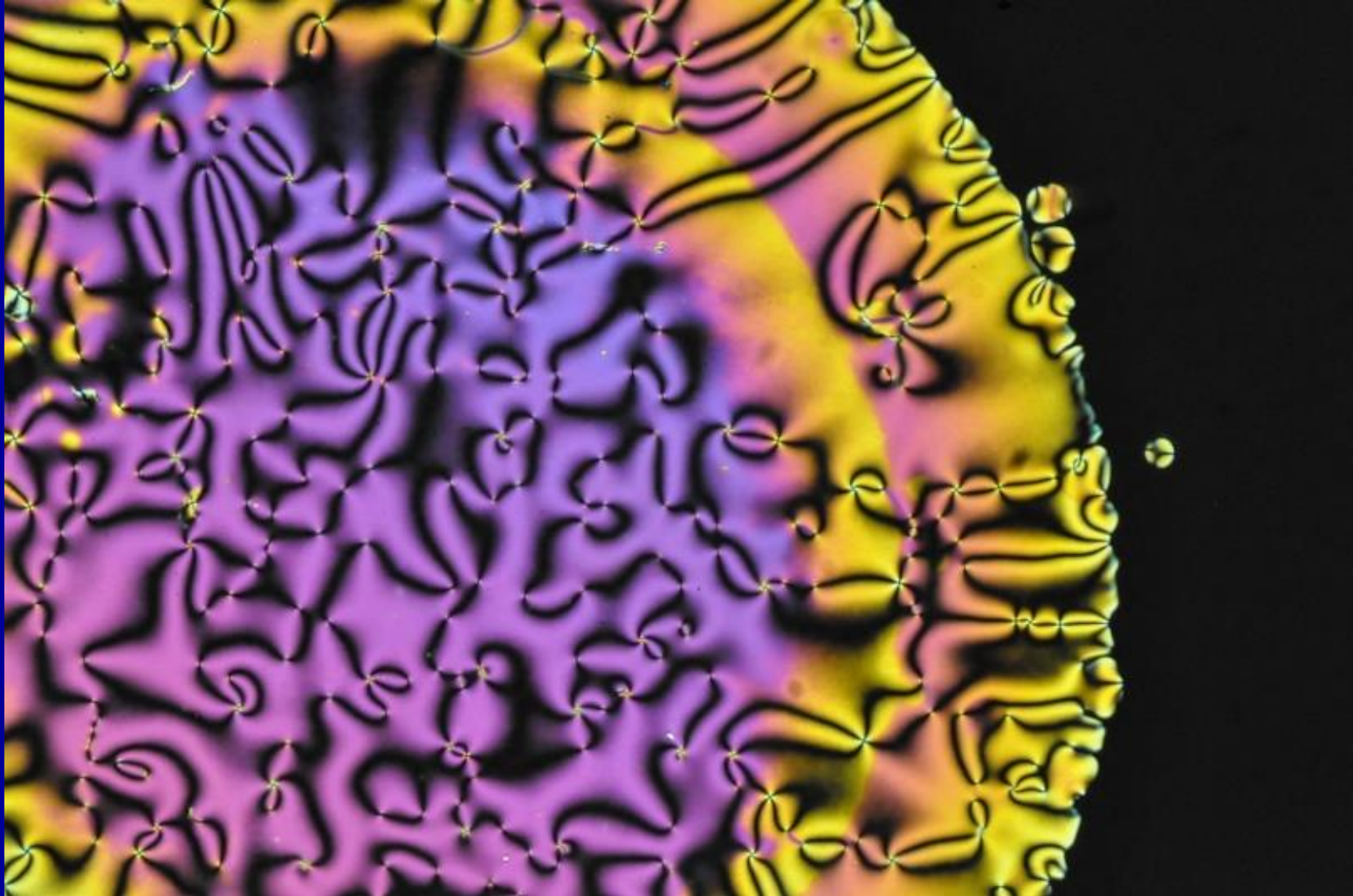
These two classes of defects can be combined according to the algebra of $\mathbb{Z}/2\mathbb{Z}$:

$$0+0=0$$

$$0+1=1$$

$$1+1=0$$

Part II. Topological defects in nematics



Nematoelasticity in a nutshell

- Consider a given director field $\mathbf{n}_0(\mathbf{r})$. Deformations about that configuration are orthogonal to $\mathbf{n}_0(\mathbf{r})$ as

$$\mathbf{n}_0 \cdot \mathbf{n}_0 = 1 \Rightarrow \mathbf{n}_0 \cdot \delta \mathbf{n} = 0$$

To simplify, if one takes $\mathbf{n}_0(\mathbf{r}) = \mathbf{e}_3$, then $\delta \mathbf{n} = (\delta n_1, \delta n_2, 0)$. Let be $\mathbf{n} = \mathbf{n}_0 + \delta \mathbf{n}$, a Taylor expansion gives:

$$\begin{pmatrix} \delta n_1 \\ \delta n_2 \end{pmatrix} = \begin{pmatrix} \frac{\partial n_1}{\partial x_1} \delta x_1 + \frac{\partial n_1}{\partial x_2} \delta x_2 + \frac{\partial n_1}{\partial x_3} \delta x_3 + \dots \\ \frac{\partial n_2}{\partial x_1} \delta x_1 + \frac{\partial n_2}{\partial x_2} \delta x_2 + \frac{\partial n_2}{\partial x_3} \delta x_3 + \dots \end{pmatrix}$$

Nematoelasticity in a nutshell

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$$\begin{pmatrix} \delta n_1 \\ \delta n_2 \end{pmatrix} = \frac{1}{2} \begin{bmatrix} \frac{\partial n_1}{\partial x_1} + \frac{\partial n_2}{\partial x_2} & 0 \\ 0 & \frac{\partial n_1}{\partial x_1} + \frac{\partial n_2}{\partial x_2} \end{bmatrix} \begin{pmatrix} \delta x_1 \\ \delta x_2 \end{pmatrix} + \frac{1}{2} \begin{bmatrix} 0 & \frac{\partial n_1}{\partial x_2} - \frac{\partial n_2}{\partial x_1} \\ -\frac{\partial n_1}{\partial x_2} + \frac{\partial n_2}{\partial x_1} & 0 \end{bmatrix} \begin{pmatrix} \delta x_1 \\ \delta x_2 \end{pmatrix} \\ + \frac{1}{2} \begin{bmatrix} \frac{\partial n_1}{\partial x_1} - \frac{\partial n_2}{\partial x_2} & \frac{\partial n_1}{\partial x_2} + \frac{\partial n_2}{\partial x_1} \\ \frac{\partial n_1}{\partial x_2} + \frac{\partial n_2}{\partial x_1} & -\frac{\partial n_1}{\partial x_1} + \frac{\partial n_2}{\partial x_2} \end{bmatrix} \begin{pmatrix} \delta x_1 \\ \delta x_2 \end{pmatrix} + \delta x_3 \cancel{\begin{pmatrix} \frac{\partial n_1}{\partial x_3} \\ \frac{\partial n_2}{\partial x_3} \end{pmatrix}}$$

Nematoelasticity in a nutshell

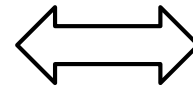
- First term:

$$\frac{1}{2} \begin{bmatrix} \frac{\partial n_1}{\partial x_1} + \frac{\partial n_2}{\partial x_2} & 0 \\ 0 & \frac{\partial n_1}{\partial x_1} + \frac{\partial n_2}{\partial x_2} \end{bmatrix} \rightarrow f_1 = \frac{1}{2} (\operatorname{div} \mathbf{n})$$

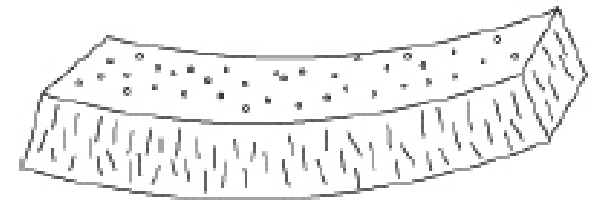
Nematoelasticity in a nutshell

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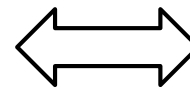
« splay »



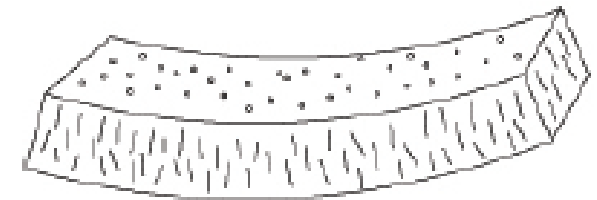
Nematoelasticity in a nutshell

- First term:

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« splay »



- Second term:

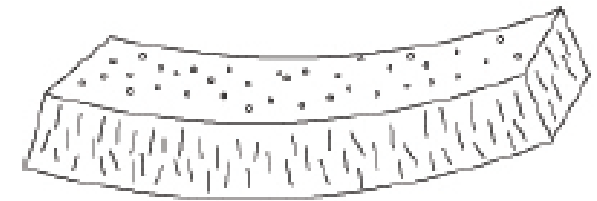
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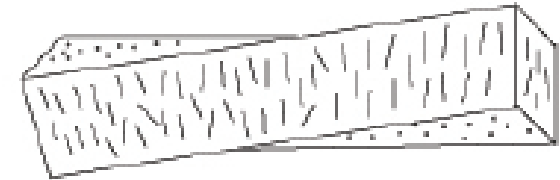
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- Second term:

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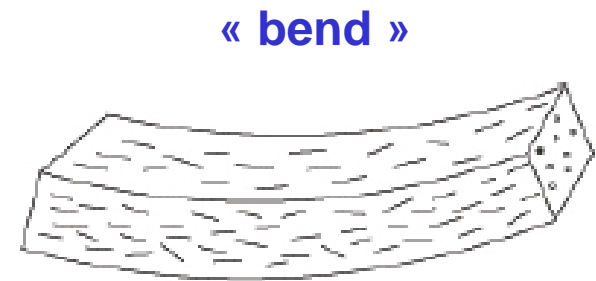
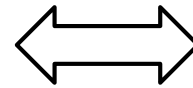
« twist »



Nematoelasticity in a nutshell

- Third term:

$$\frac{1}{2} \begin{bmatrix} \frac{\partial n_1}{\partial x_1} - \frac{\partial n_2}{\partial x_2} & \frac{\partial n_1}{\partial x_2} + \frac{\partial n_2}{\partial x_1} \\ \frac{\partial n_1}{\partial x_2} + \frac{\partial n_2}{\partial x_1} & -\frac{\partial n_1}{\partial x_1} + \frac{\partial n_2}{\partial x_2} \end{bmatrix} \rightarrow f_3 \approx \frac{1}{2} (\mathbf{n} \wedge \mathbf{curl} \mathbf{n})$$

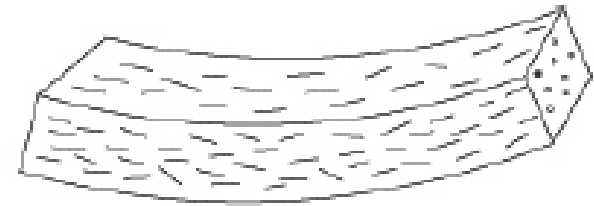


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« bend »



- Similarwise to the harmonic oscillator, the (simplest) Frank-Oseen free energy describing nematoelasticity writes as

$$F = \frac{1}{2} K X^2 \simeq \frac{1}{2} K_1 (\text{div} \mathbf{n})^2 + \frac{1}{2} K_2 (\mathbf{n} \cdot \mathbf{curl} \mathbf{n})^2 + \frac{1}{2} K_3 (\mathbf{n} \wedge \mathbf{curl} \mathbf{n})^2$$

Planar distortions

- “One constant approximation” (isotropic elasticity) : $K_1 \sim K_2 \sim K_3 = K \sim \frac{E_0}{L}$

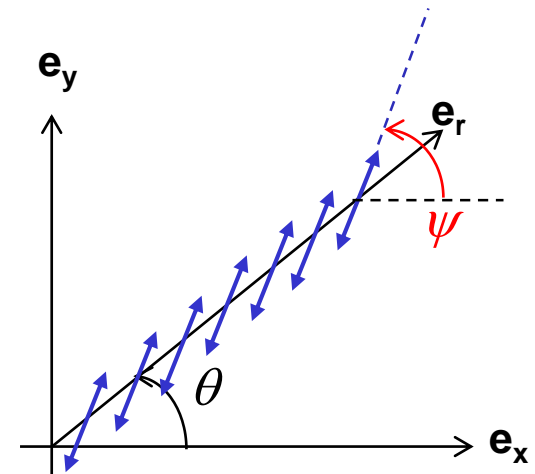
$$\Rightarrow F = \frac{K}{2} \left[(\operatorname{div} \mathbf{n})^2 + (\mathbf{n} \cdot \operatorname{curl} \mathbf{n})^2 + (\mathbf{n} \wedge \operatorname{curl} \mathbf{n})^2 \right] = \frac{K}{2} \left[(\operatorname{div} \mathbf{n})^2 + (\operatorname{curl} \mathbf{n})^2 \right]$$

- For planar configurations of the director field (x-y), tedious calculations give

$$\mathbf{n}(r, \theta) = \begin{pmatrix} \cos \psi(r, \theta) \\ \sin \psi(r, \theta) \\ 0 \end{pmatrix} \Rightarrow F = \frac{K}{2} (\operatorname{grad} \psi)^2$$

Euler-Lagrange equation for the Frank-Oseen energy then reduces to

$$\Delta \psi = \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial \psi}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2 \psi}{\partial \theta^2} = 0$$



Planar distortions

- One seeks simple solutions that are not depending on r : $\frac{d^2\psi}{d\theta^2} = 0 \Rightarrow \psi(\theta) = m\theta + \psi_0$

\Rightarrow After a full turn about the z-axis, $\oint_{\theta=2\pi} d\psi = 2\pi m$

m = winding number (in \mathbb{R})

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- Constraints on k : the direction of \mathbf{n} is well-defined at each point $\Rightarrow \oint_{\theta=2\pi} d\psi = 2\pi m = k\pi \quad k \in \mathbb{Z}$

(\mathbb{Z}_2 symmetry of the nematic state)

$$\Rightarrow m = \pm\frac{1}{2}, \pm 1, \pm\frac{3}{2}, \pm 2, \dots$$

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Whole-numbered disclinations

► What does a distorted nematic locally look like ?

It depends on the parameters (m, ψ_0) . Let us learn a little more from several examples:

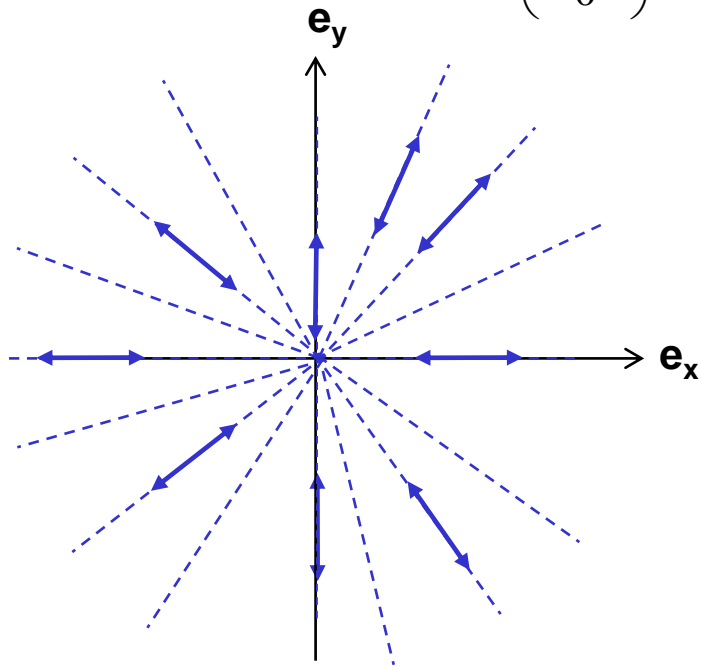
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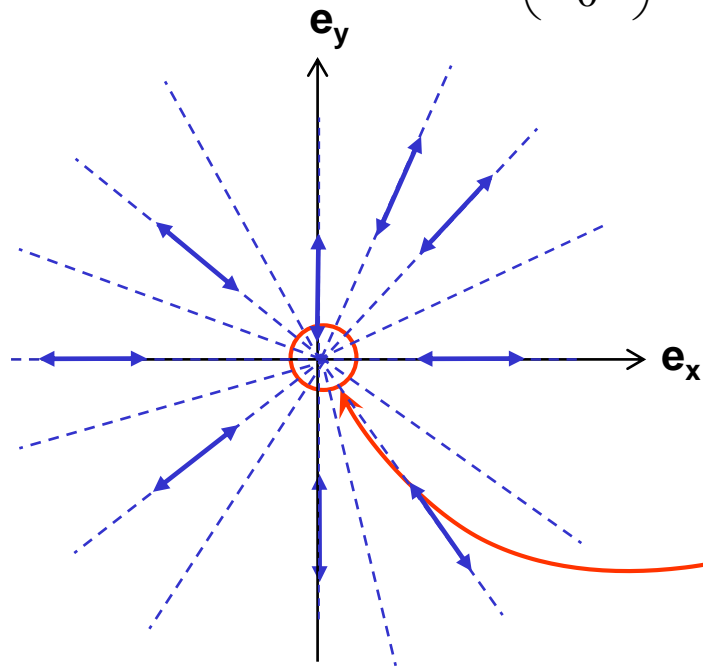


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*singularity of orientation along this line
= linear « defect »
= « disclination line »*

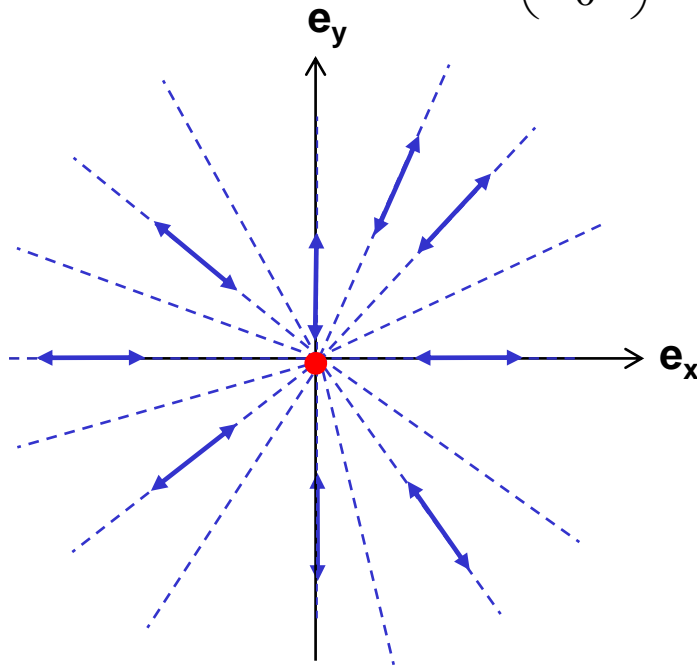
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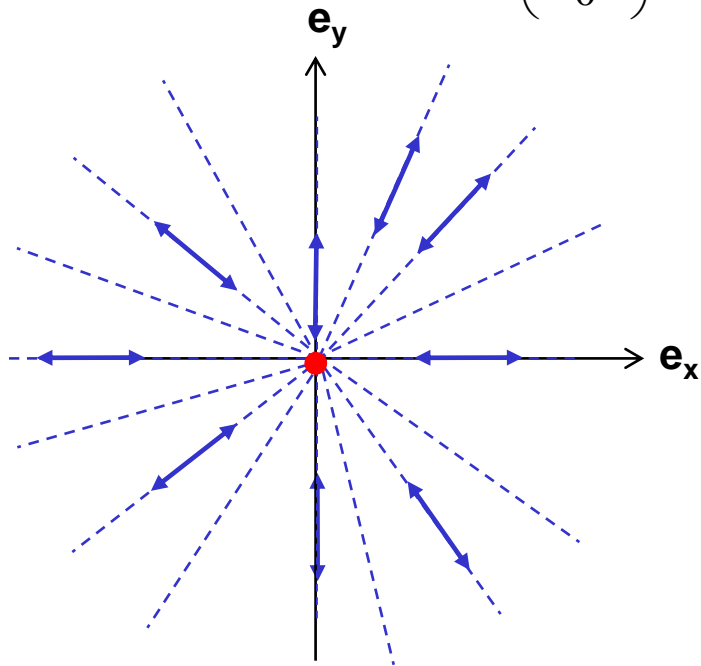


Whole-numbered disclinations

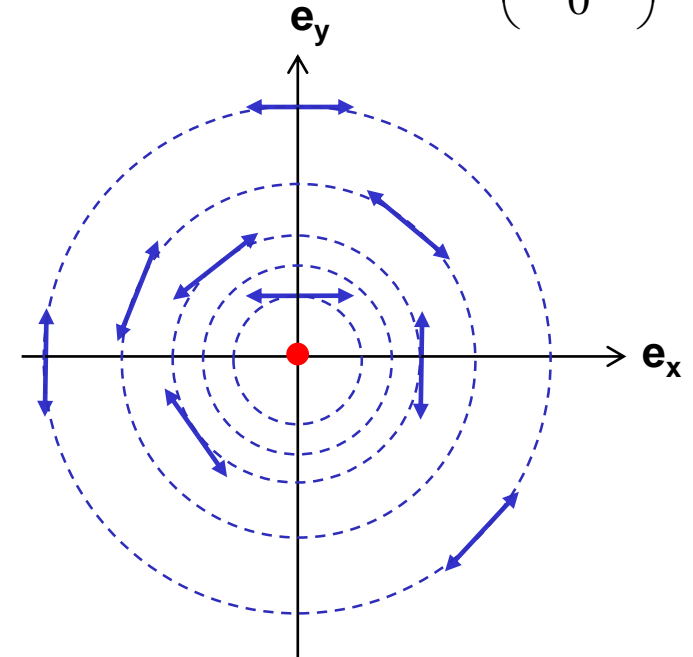
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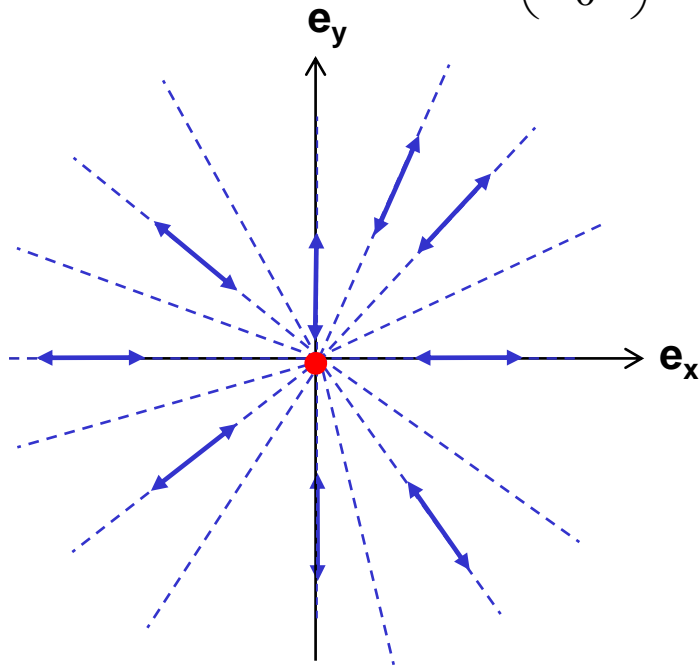


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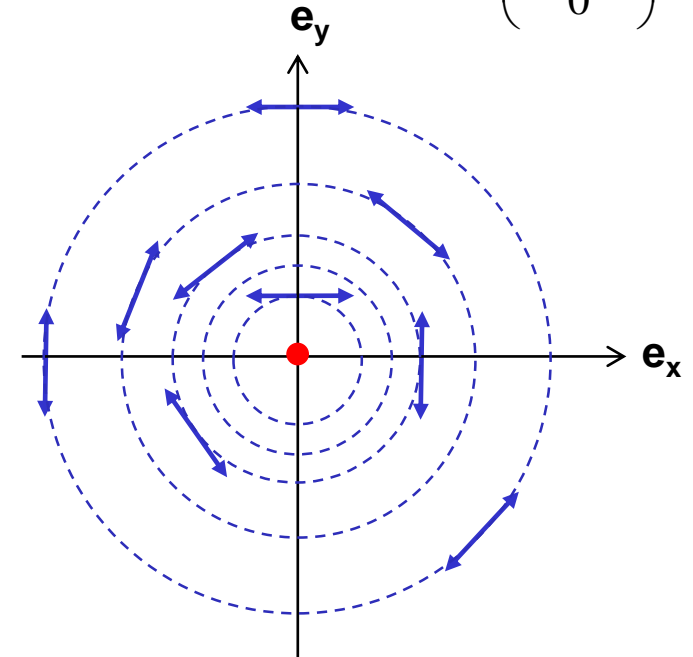
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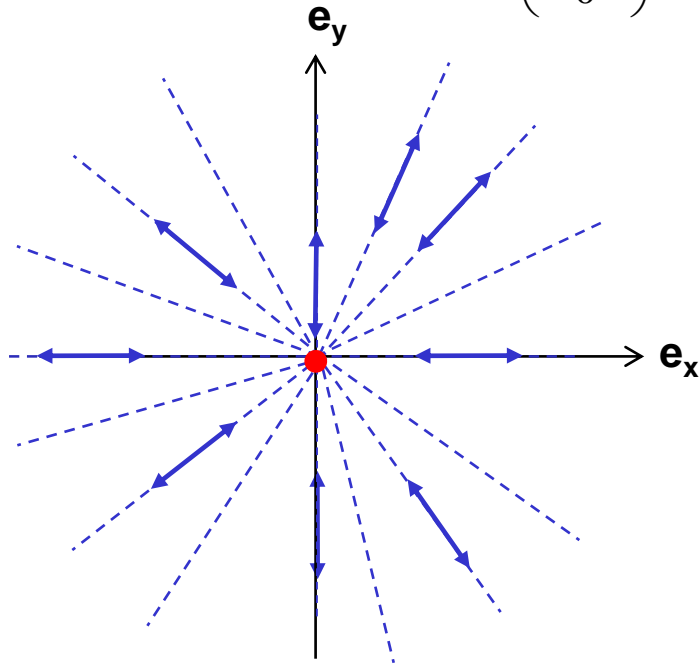


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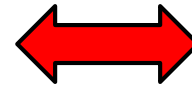
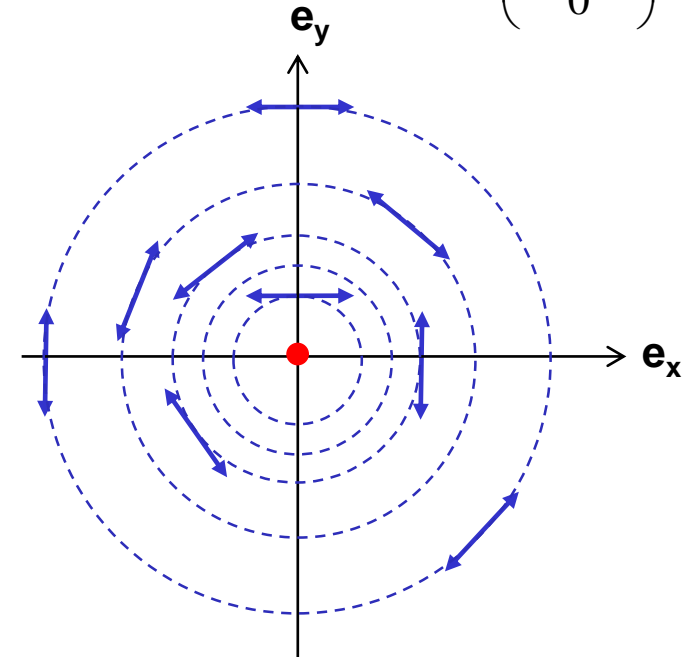
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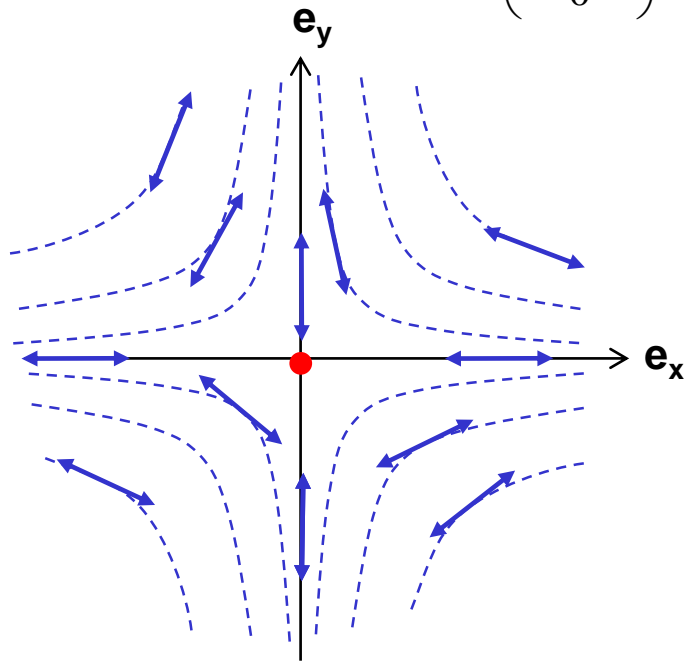
**Topologically
 equivalent
 ($D \geq 2$)**

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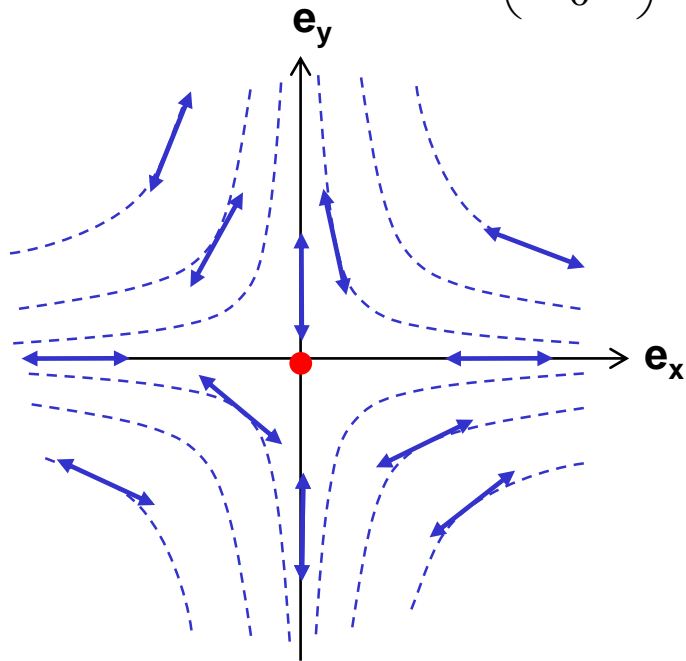


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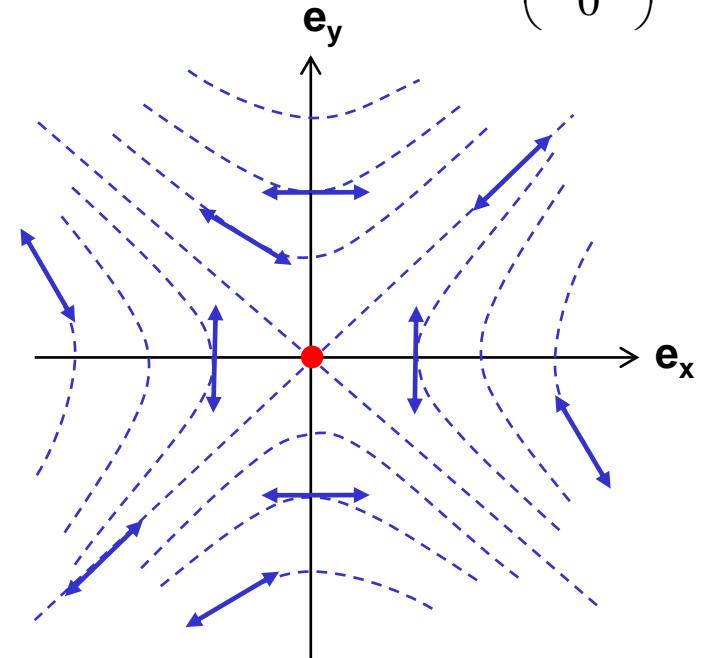
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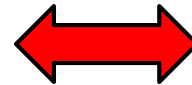
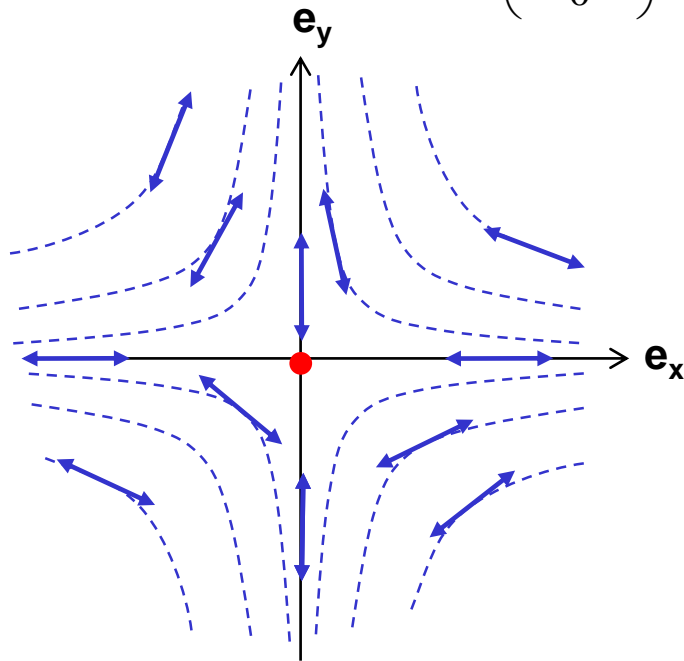


Whole-numbered disclinations

► What does a distorted nematic locally look like ?

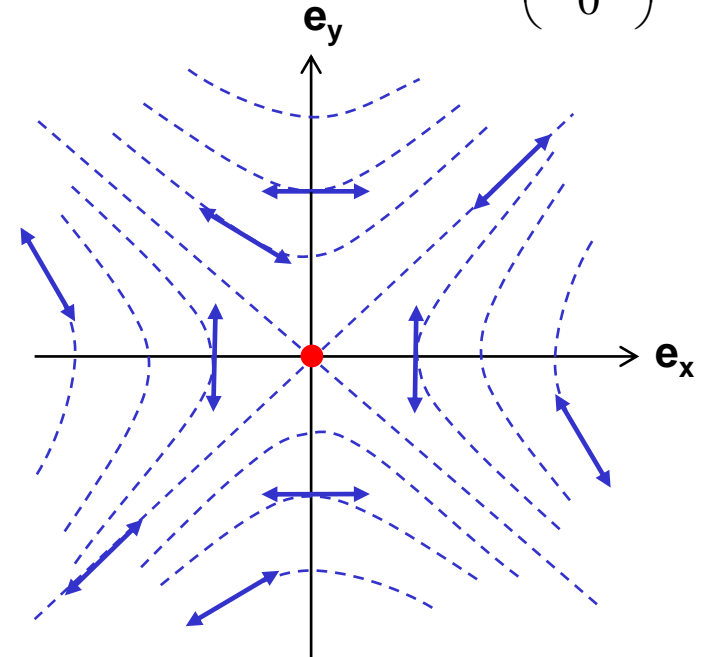
It depends on the parameters (m, ψ_0) . Let us learn a little more from several examples:

$$m = -1, \psi_0 = 0 \Rightarrow \mathbf{n}(\theta) = \begin{pmatrix} \cos \theta \\ -\sin \theta \\ 0 \end{pmatrix}$$



**Topologically
 equivalent
 ($D \geq 2$)**

$$m = -1, \psi_0 = \pi / 2 \Rightarrow \mathbf{n}(\theta) = \begin{pmatrix} \sin \theta \\ \cos \theta \\ 0 \end{pmatrix}$$

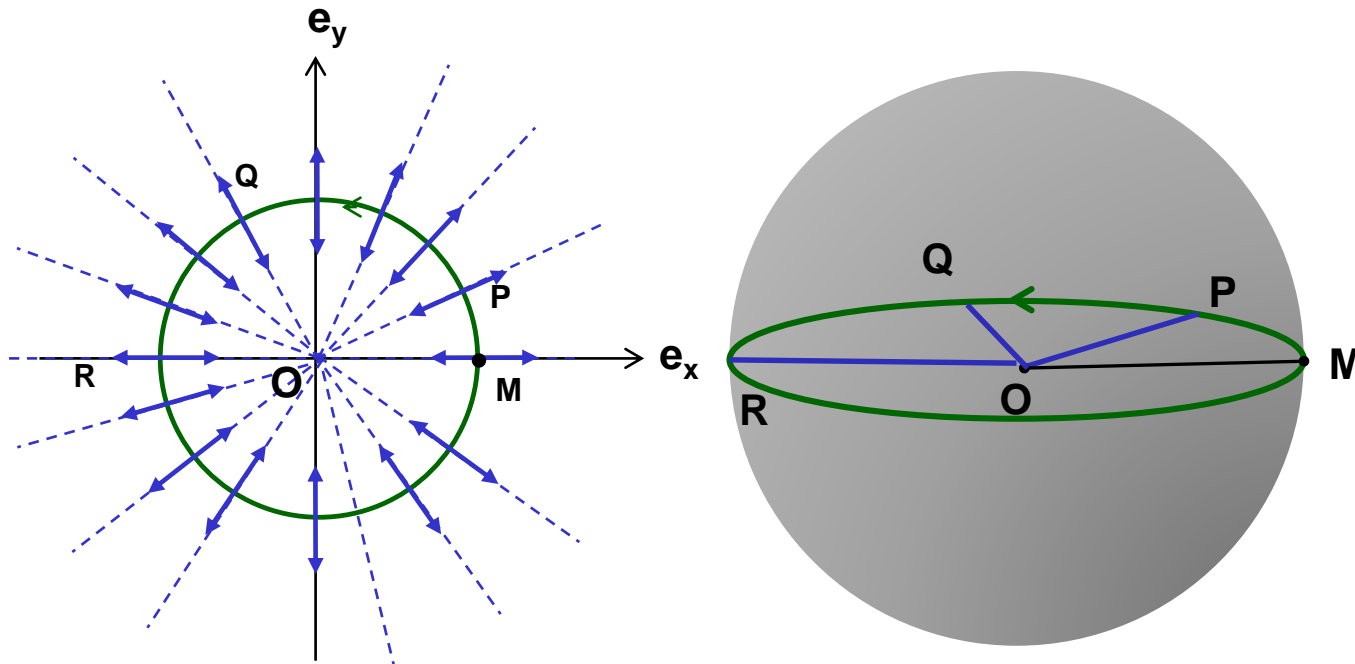


Whole-numbered disclinations

- ▶ Is m the label of π_1 equivalence classes ?

Whole-numbered disclinations

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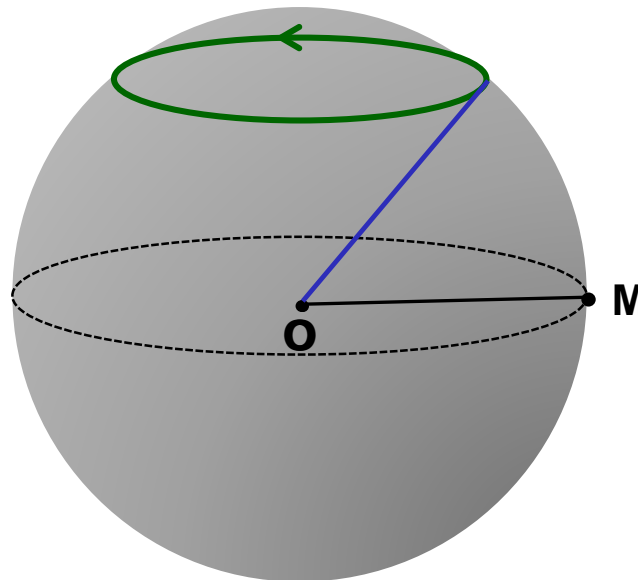
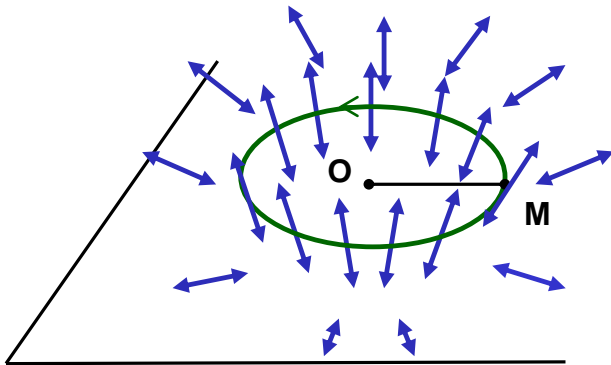


$$\pi_1(\mathbb{R}P^2) = \mathbb{Z}_2 = \{0, 1\}$$

Whole-numbered disclinations

► Is m the label of π_1 equivalence classes ?

The « magic trick » = escape in the third dimension !

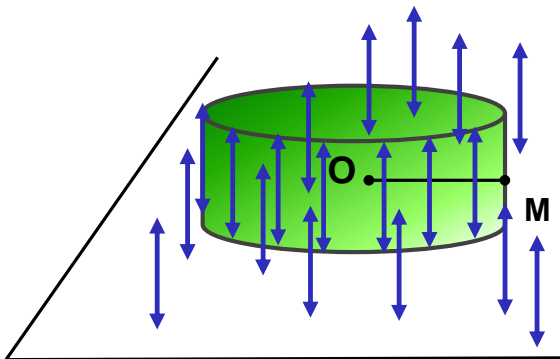


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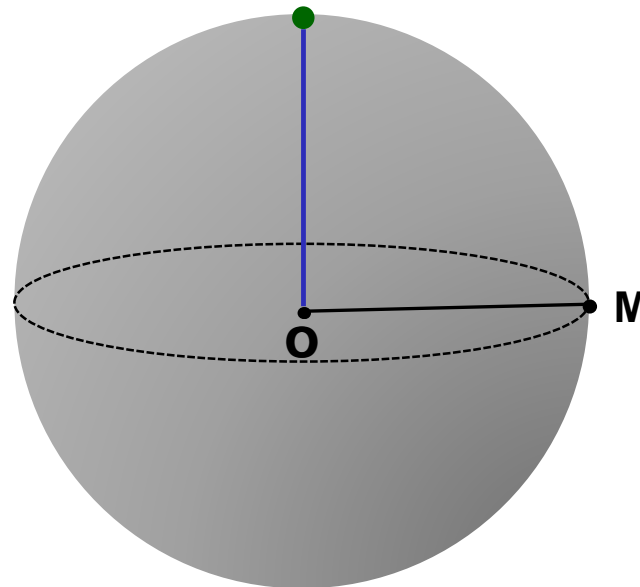
Whole-numbered disclinations

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The « magic trick » = escape in the third dimension !



Regular ribbon



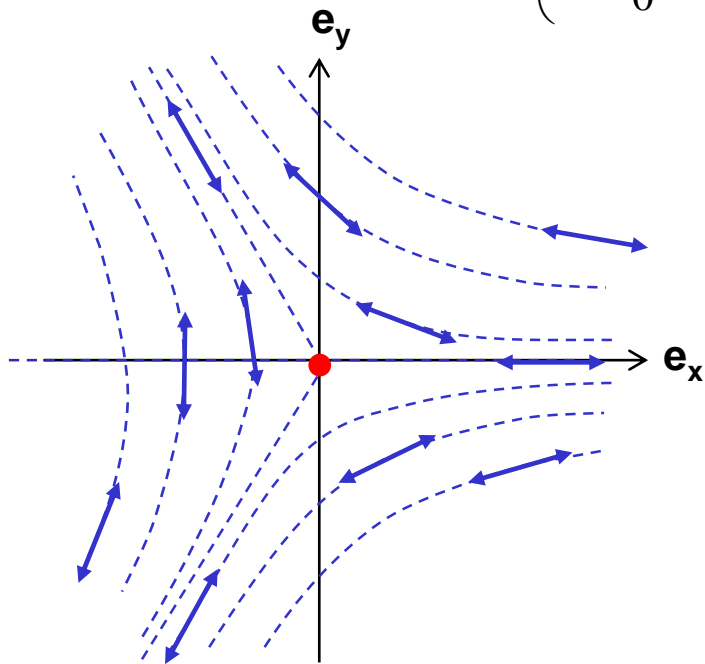
$$\pi_1(\mathbb{R}P^2) = \mathbb{Z}_2 = \{0, 1\}$$

The homotopy loop can be shrunk to a point = **topologically removable defect**

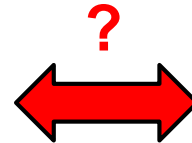
Likewise for $m=-1 \Rightarrow$ Disclinations of integer strengths belong to the trivial homotopy class $N=0$.

Moebius disclinations

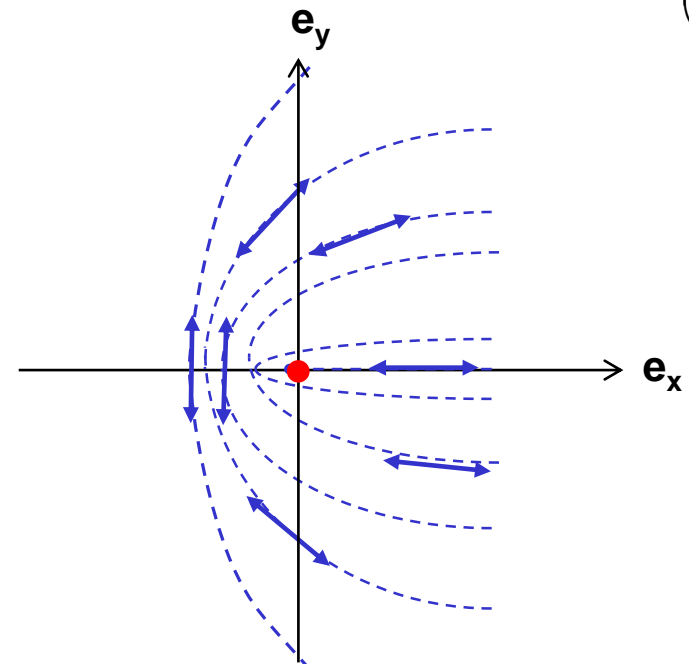
$$m = -1/2, \psi_0 = 0 \Rightarrow \mathbf{n}(\theta) = \begin{pmatrix} \cos(\theta/2) \\ -\sin(\theta/2) \\ 0 \end{pmatrix}$$



« Trefoil »

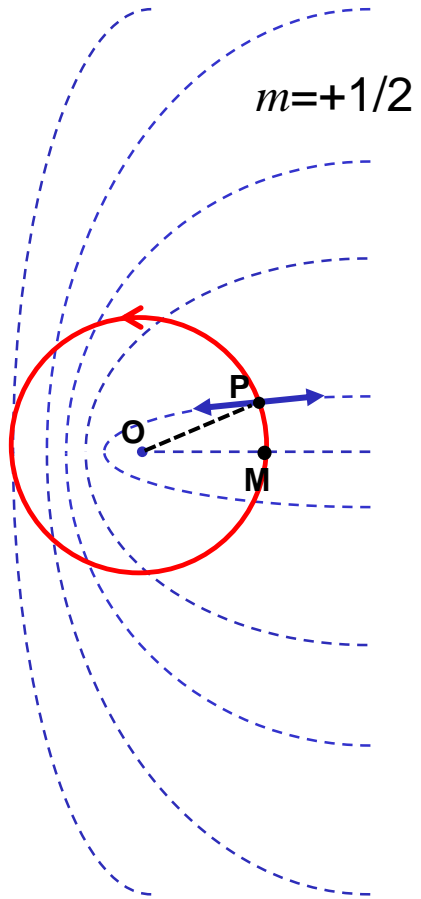


$$m = +1/2, \psi_0 = 0 \Rightarrow \mathbf{n}(\theta) = \begin{pmatrix} \cos(\theta/2) \\ \sin(\theta/2) \\ 0 \end{pmatrix}$$

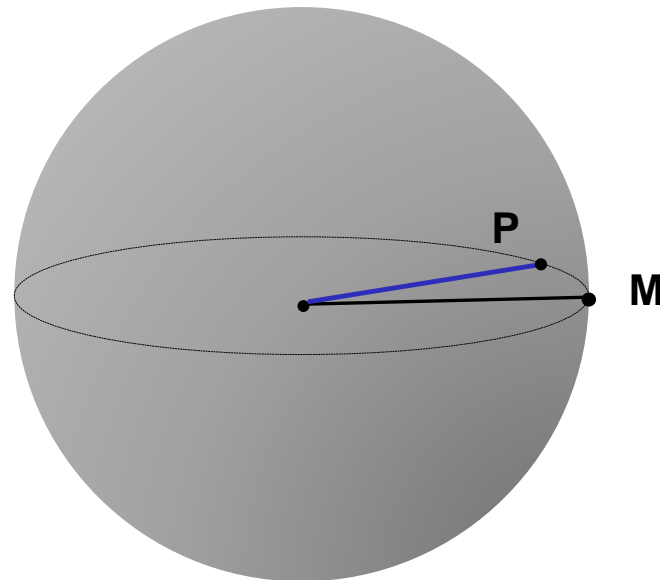


« Comet »

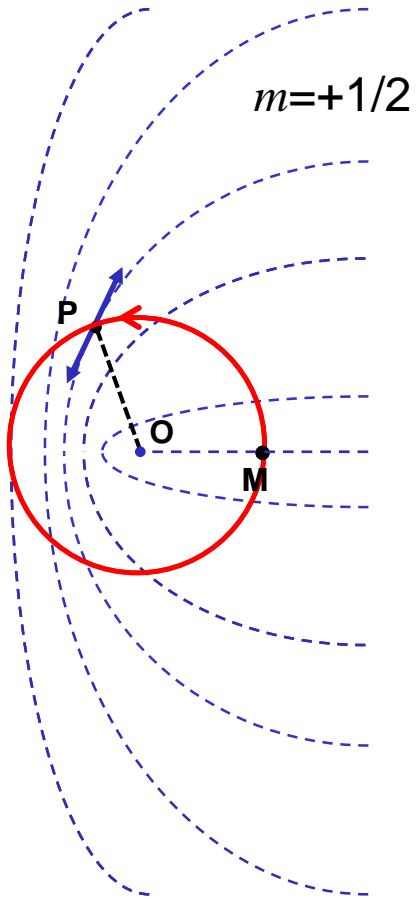
Moebius disclinations



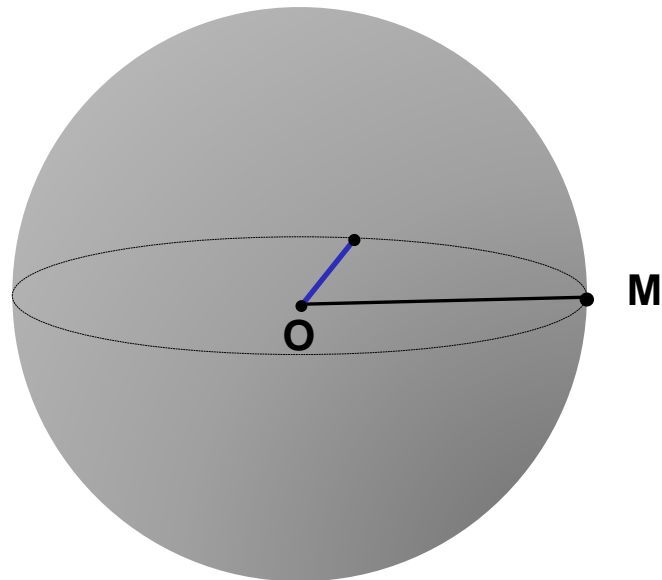
$$\pi_1(\mathbb{R}P^2) = \mathbb{Z}_2 = \{0, 1\}$$



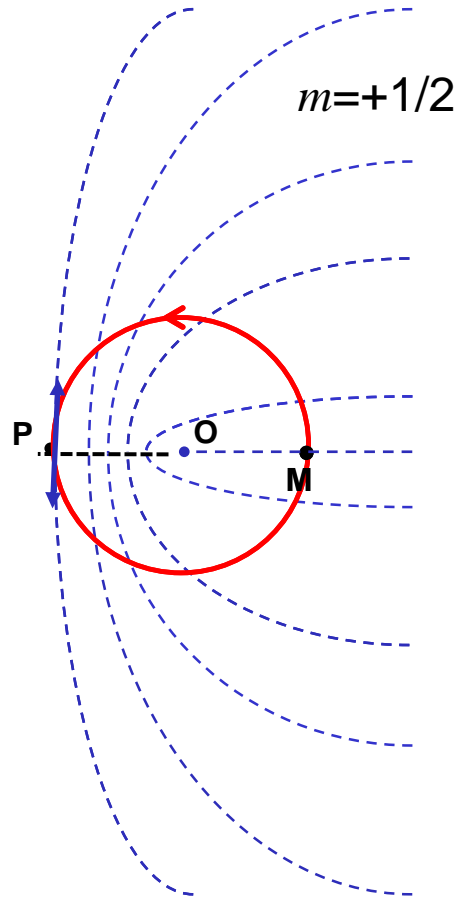
Moebius disclinations



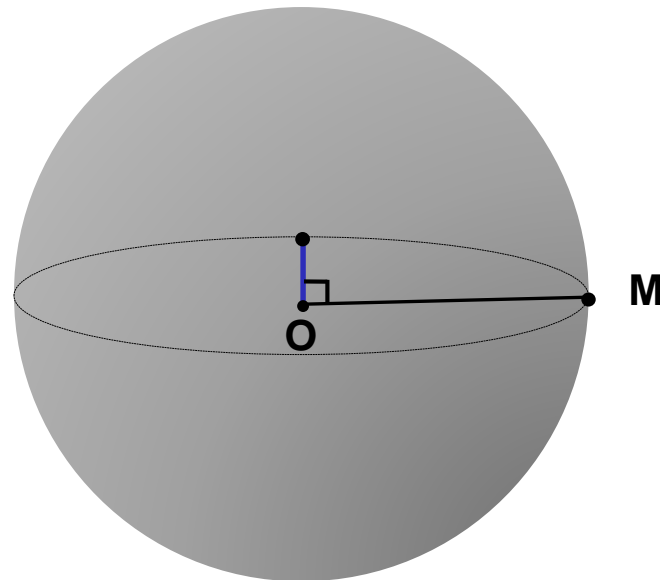
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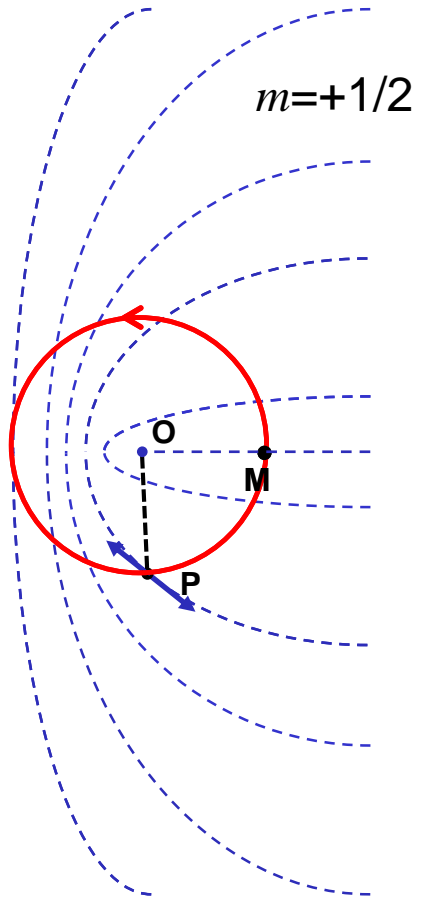
Moebius disclinations



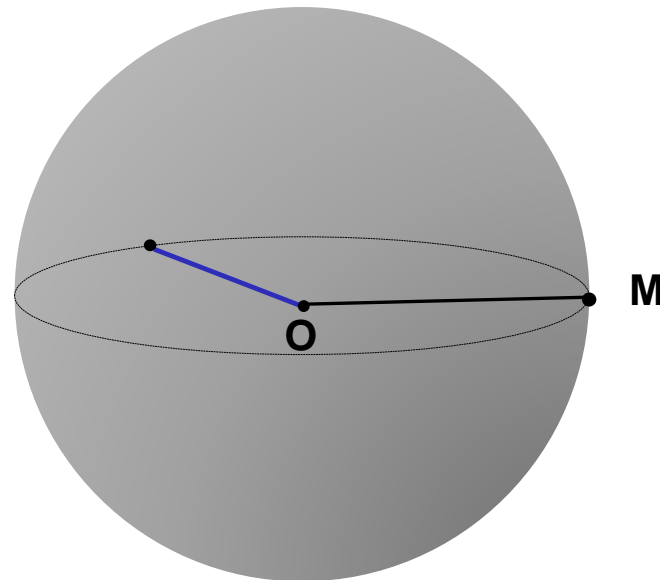
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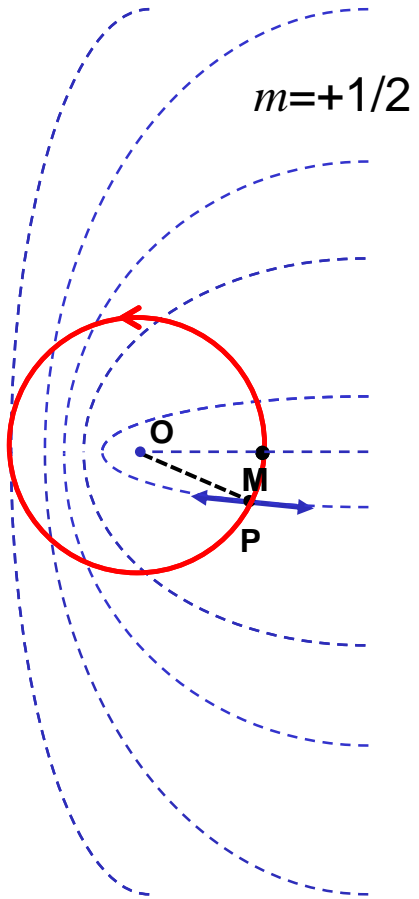
Moebius disclinations



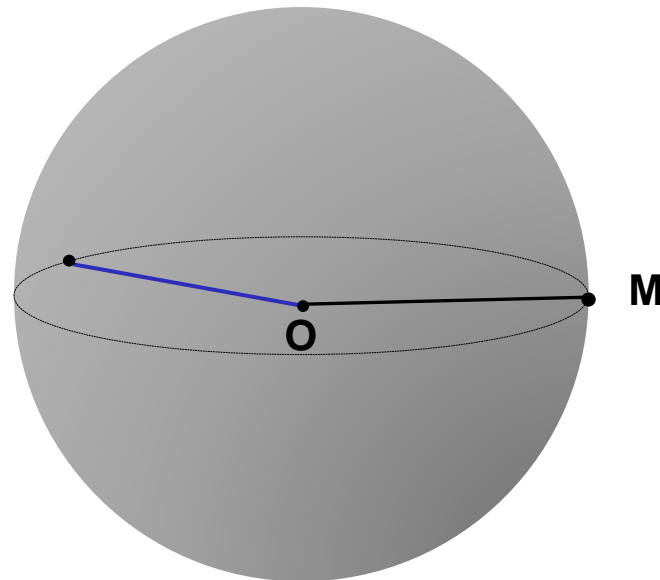
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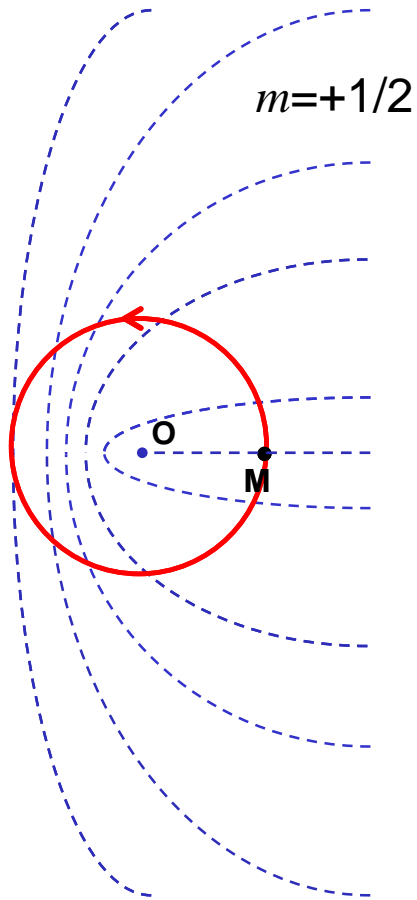
Moebius disclinations



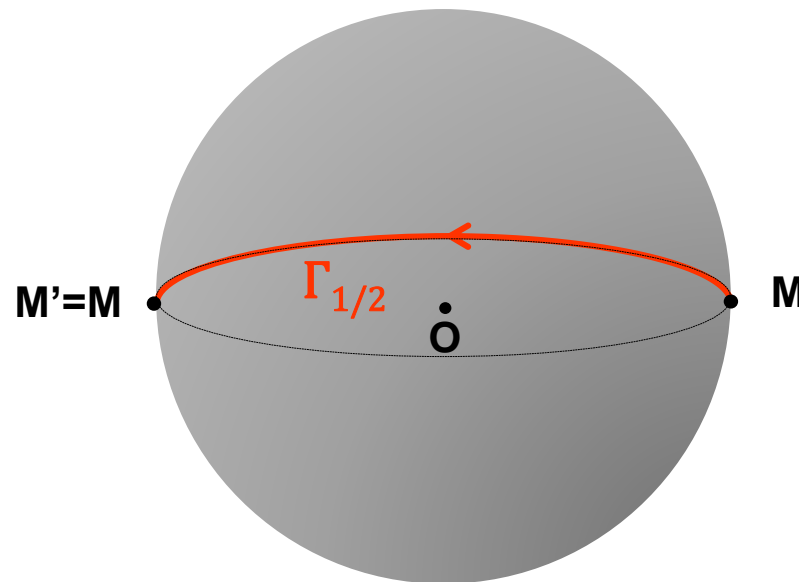
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Moebius disclinations



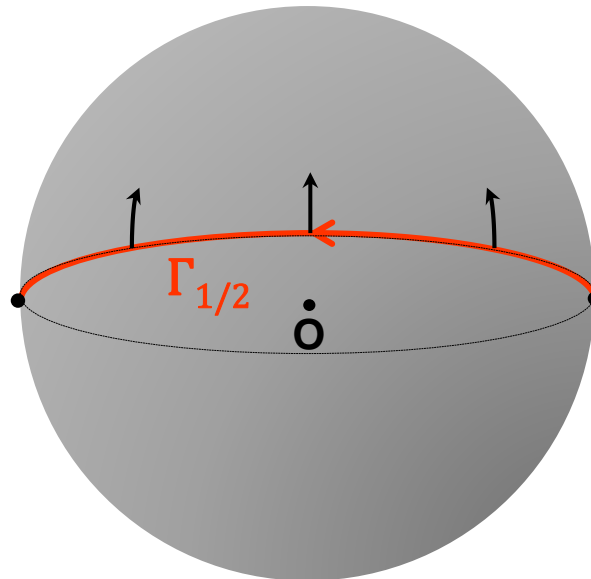
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The homotopy loop cannot be shrunk to a point
= no escape in the 3rd dimension = **topologically non-removable defect**.

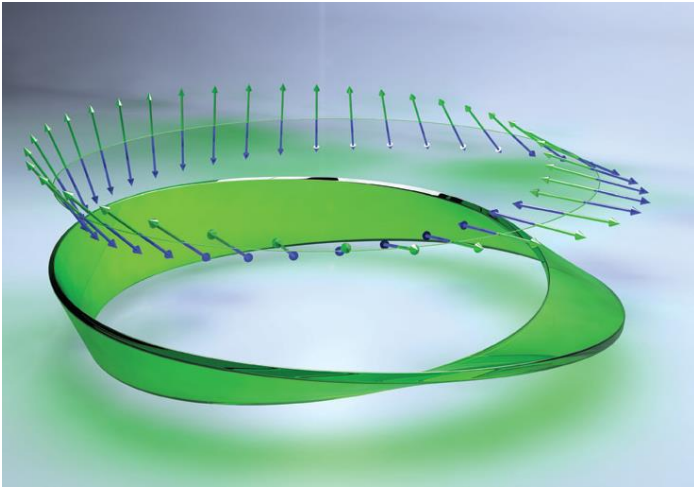
Moebius disclinations

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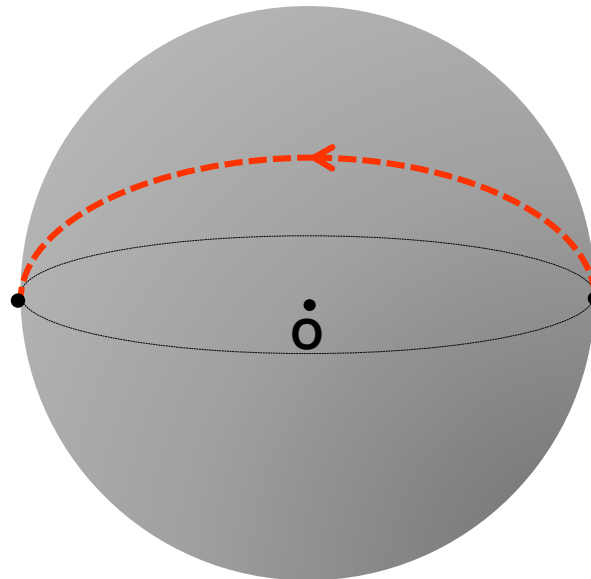


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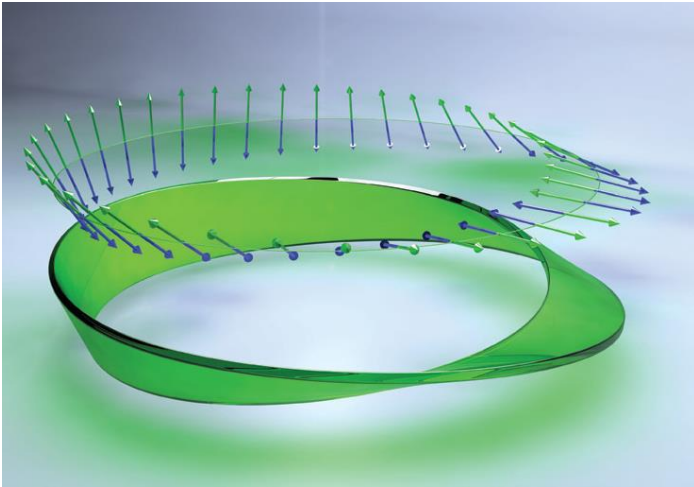


Moebius ribbon

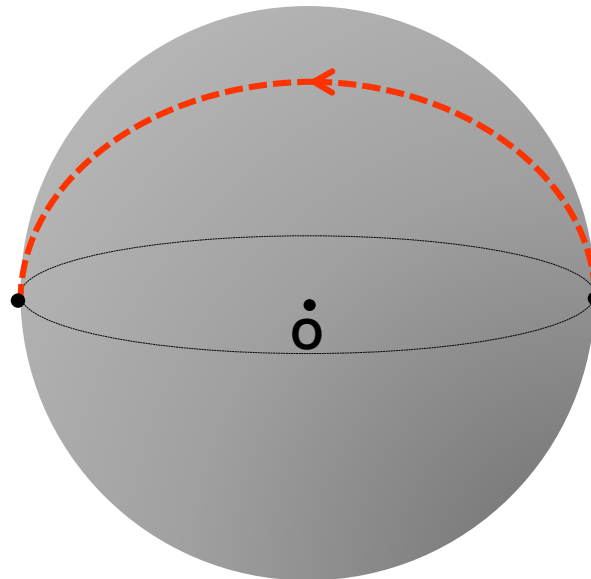


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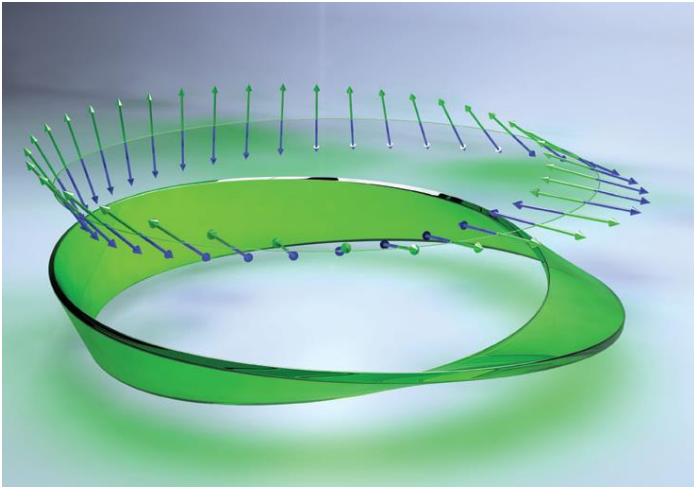


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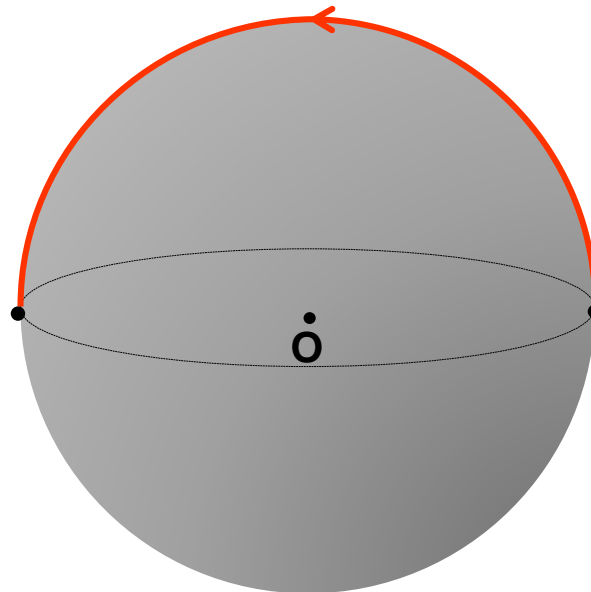


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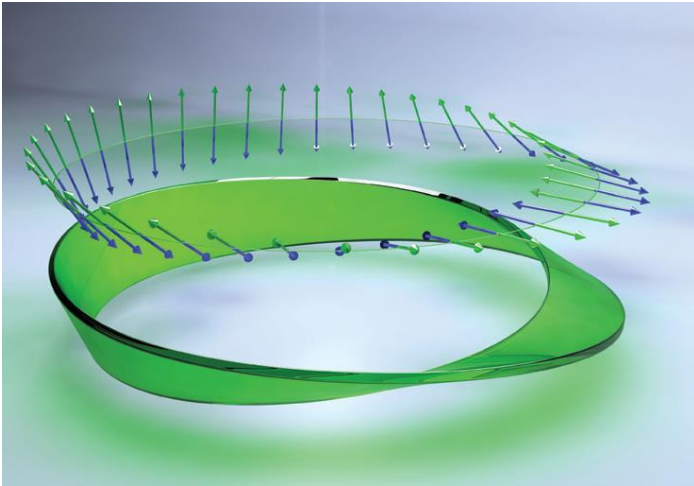


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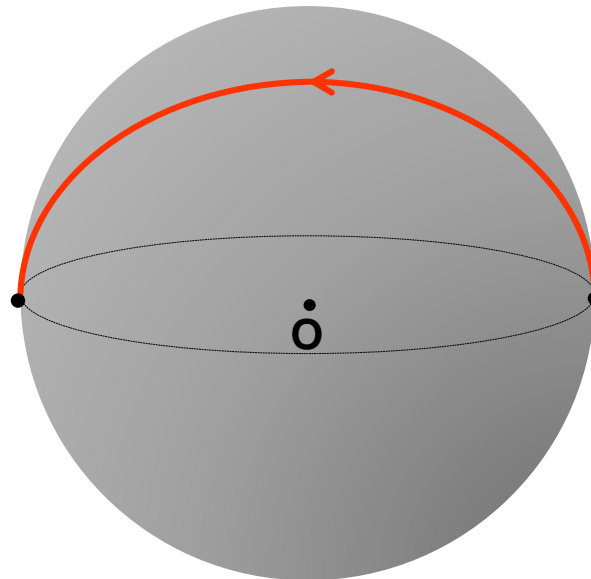


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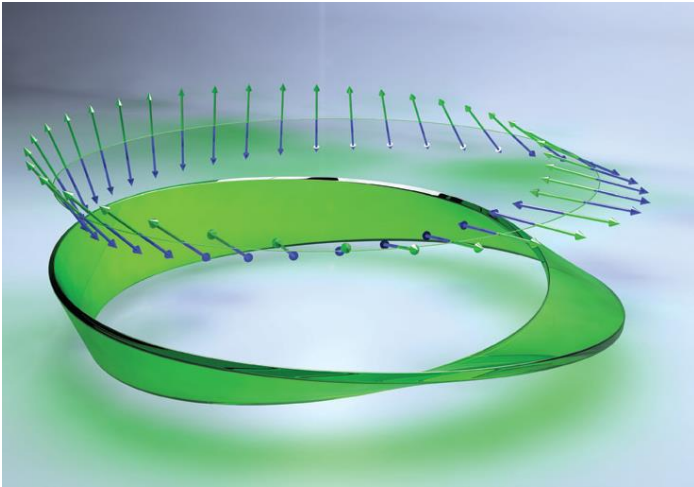


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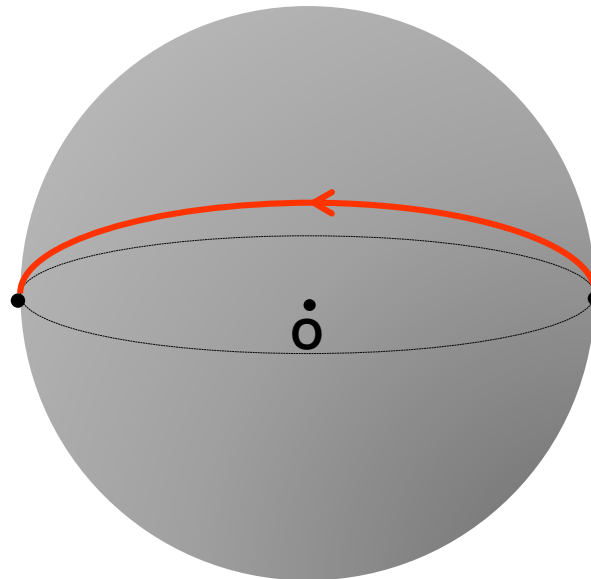


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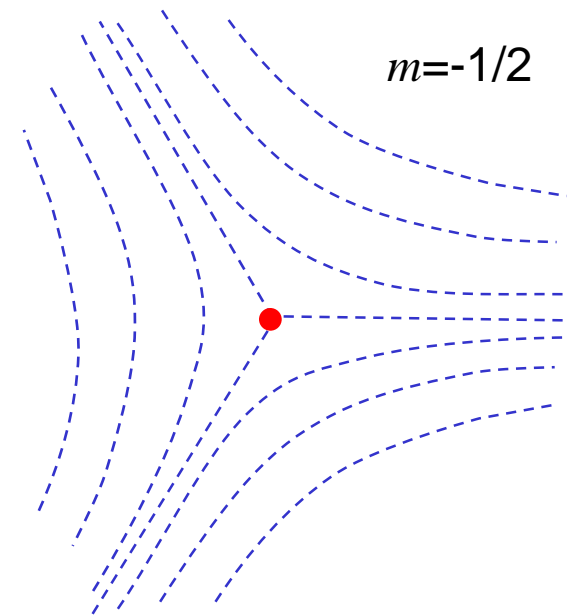
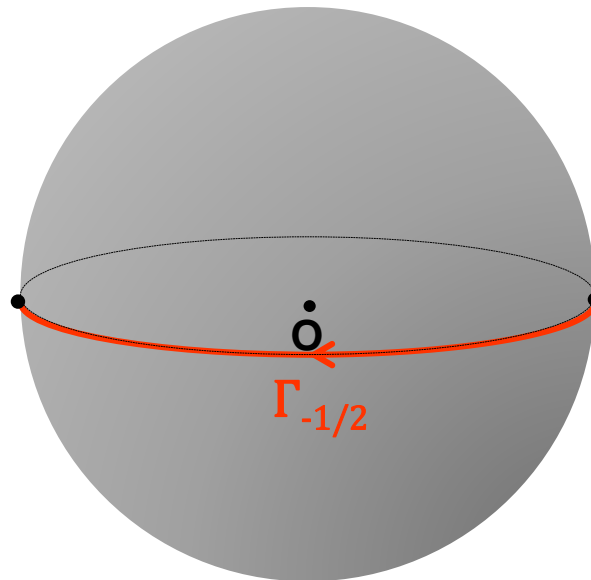
Moebius ribbon



Moebius disclinations

AA Balinskii, GE Volovik, EI Kats. Sov. Phys. JETP 60 (1984)

$$\pi_1(\mathbb{R}P^2) = \mathbb{Z}_2 = \{0, 1\}$$



$\Gamma_{1/2}$ and $\Gamma_{-1/2}$ are topologically equivalent.

⇒ Disclinations of half-integer strengths belong to the same homotopy class $N=1$

Solution to the riddle

⇒ Although m is called the « topological charge » of the defect, it is the absolute value of m that matters for topology.

$$N = 1 - E(|m|)$$

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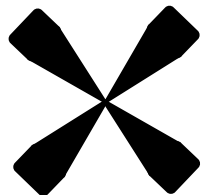
$$N = 1 - E(|m|)$$

- Polarising microscopy reveals « Schlieren patterns », which depend on this topological invariant, as the number of dark brushes = $4 \times |m|$.

$$m = \pm 1/2$$



$$m = \pm 1$$

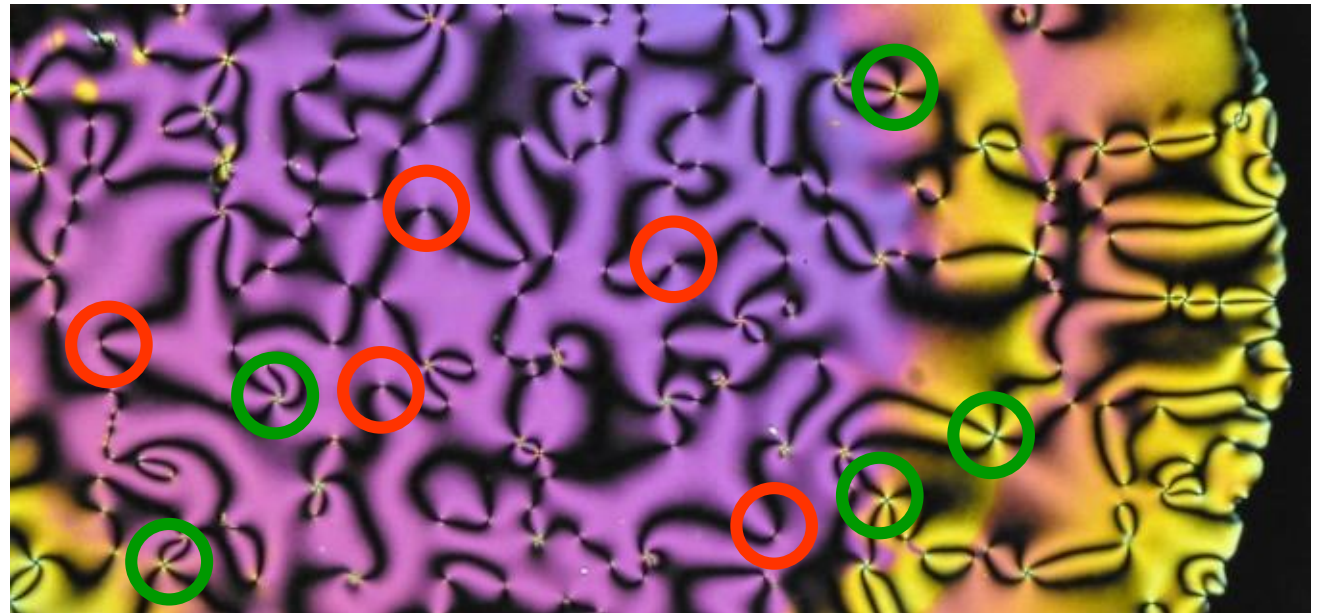
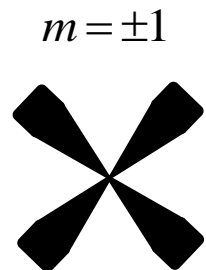
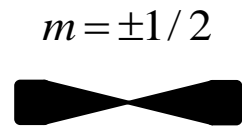


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Algebra of linear defects

S Chandrasekhar. Liquid crystals (1984)

- The set is the quotient group $\mathbb{Z} / 2\mathbb{Z} = \{0, 1\}$ with a law of addition $+$.

$$0+0=0$$

$$0+1=1$$

$$1+1=0$$

Analogously to electrostatics, defects of opposite charges attract each other (repell otherwise):

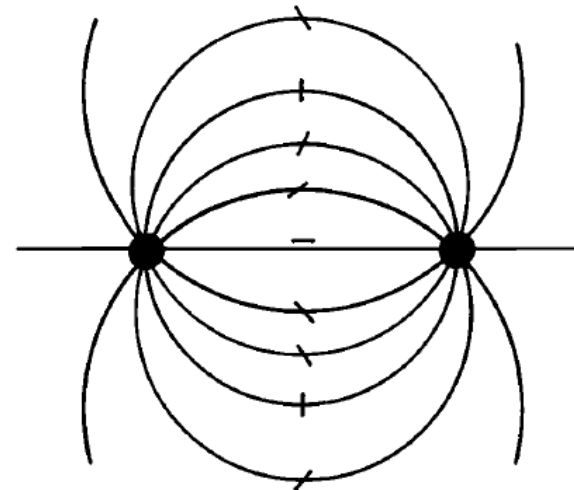


Fig. 3.5.3. Curves of equal alignment around a pair of singularities of equal and opposite strengths. The orientations marked on the circles refer to the case $s = 1$, $c = 0$.

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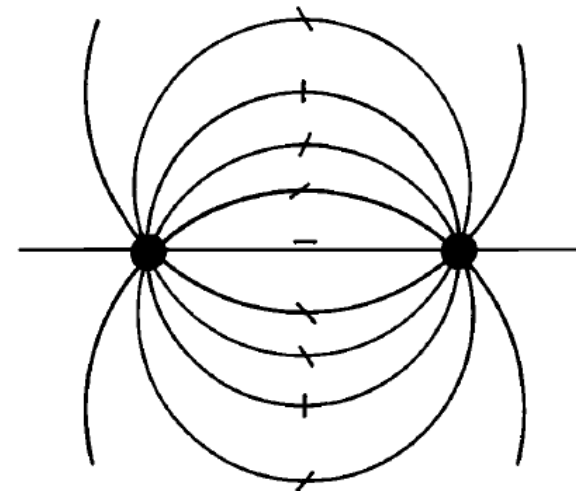


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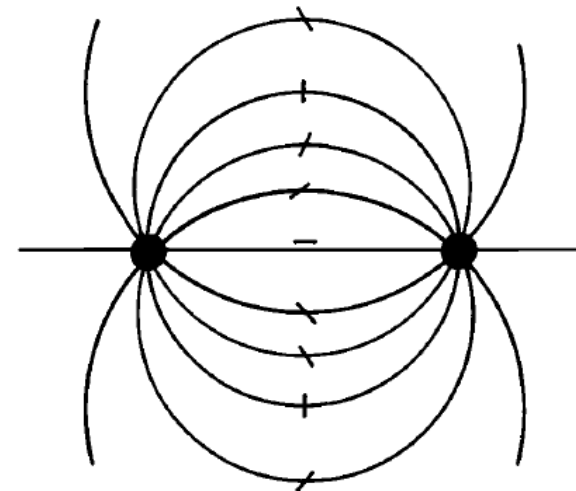


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Algebra of linear defects

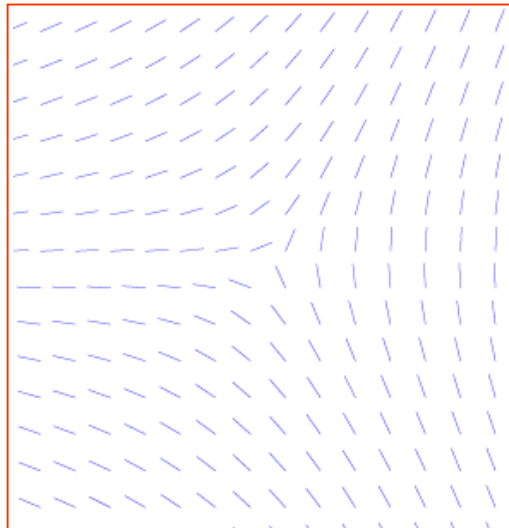
C Zhang. PhD thesis. Carnegie Mellon university (2017)

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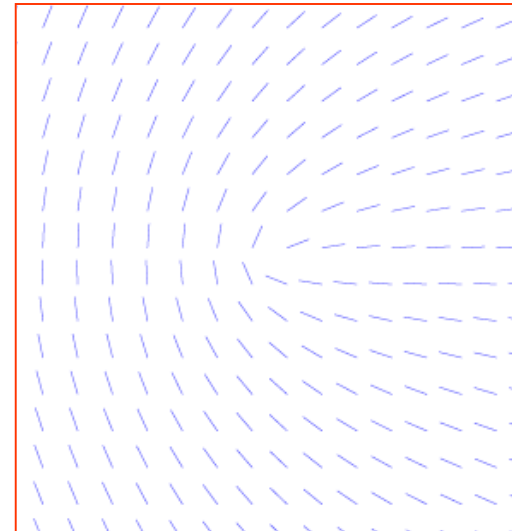
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$m=-1/2$



$m=+1/2$

Algebra of linear defects

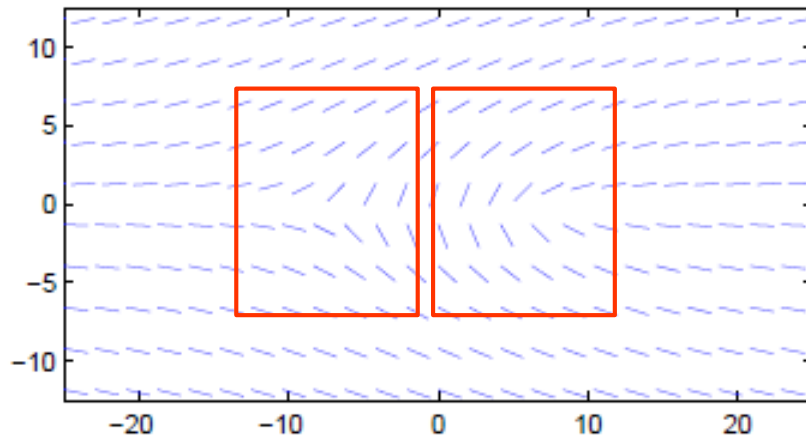
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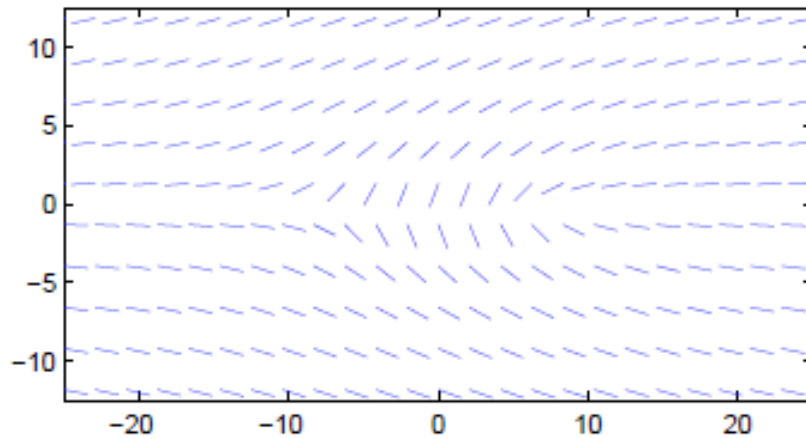
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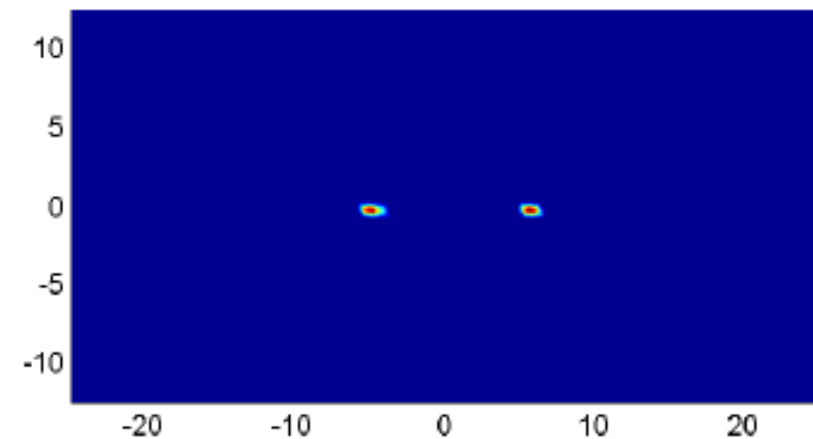
$$0+0=0$$

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(a) Director snapshot at t=0.



(b) Energy density plot at t=0.

Algebra of linear defects

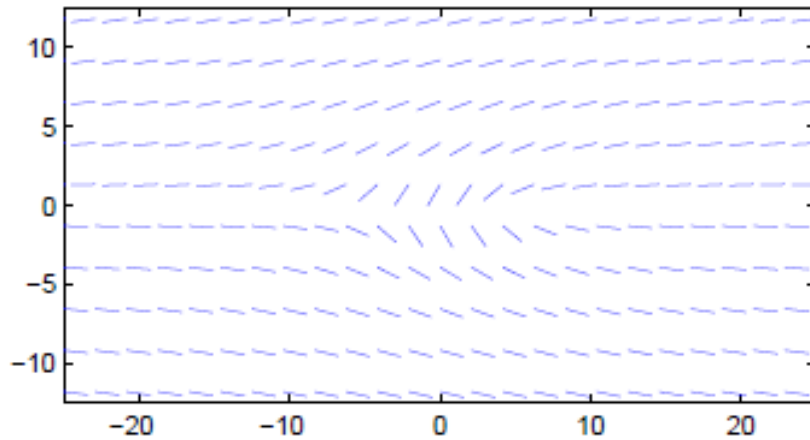
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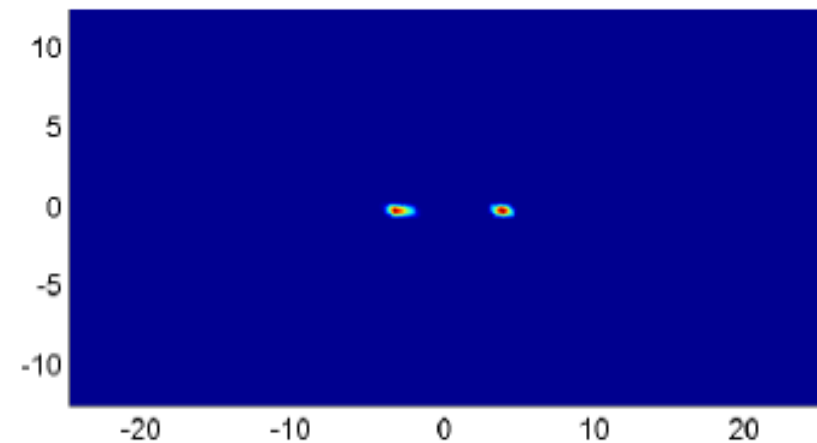
$$0+0=0$$

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(c) Director snapshot at t=1.



(d) Energy density plot at t=1.

Algebra of linear defects

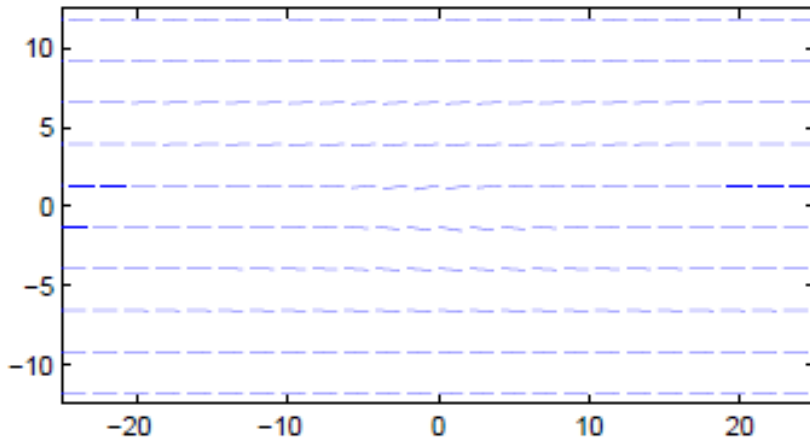
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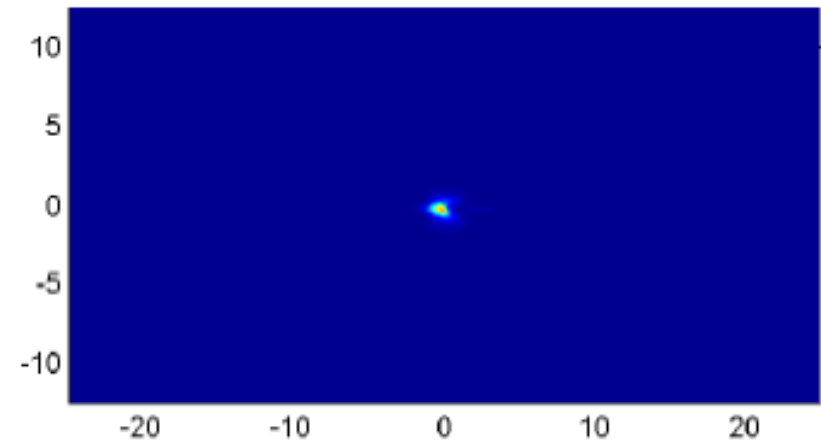
$$0+0=0$$

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(e) Director snapshot at $t=1.2$.



(f) Energy density plot at $t=1.2$.

Algebra of linear defects

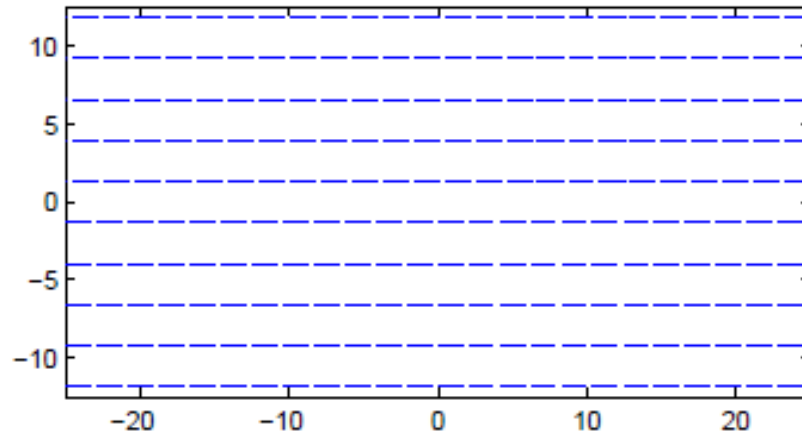
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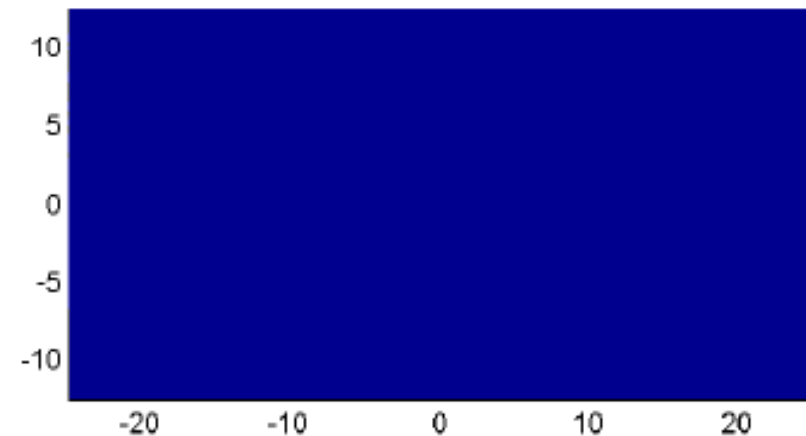
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(g) Director snapshot at $t=1.25$.

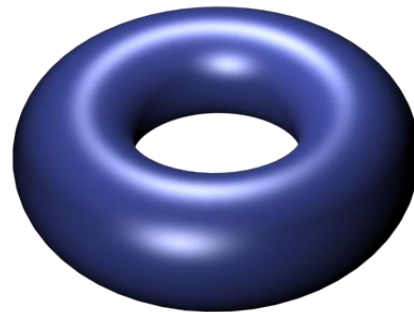


(h) Energy density plot at $t=1.25$.

Other topological numbers

► Is $|m|$ enough to characterize the topology of a line defect ?

Locally yes, but globally no, as a line defect can self-connect, entangle with itself (« nematic braids ») or more...

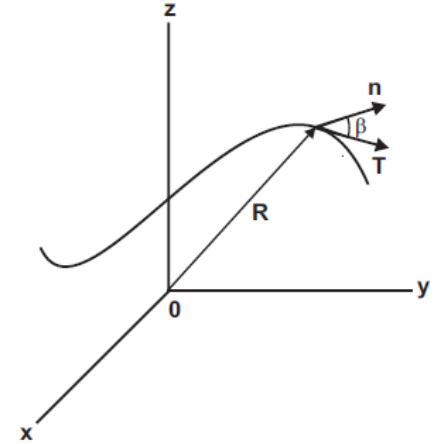


To go further: self-linking number, Jänich's index, Poincaré-Hopf's index...

Fermat-Grandjean principle (1919)

- Extraordinary light paths obey a least action principle

$$\delta \left[\int_A^B N_e(\mathbf{r}) dl \right] = 0 = \delta \left[\int_A^B \sqrt{\varepsilon_{\perp} \cos^2 \beta(\mathbf{r}) + \varepsilon_{\parallel} \sin^2 \beta(\mathbf{r})} dl \right]$$

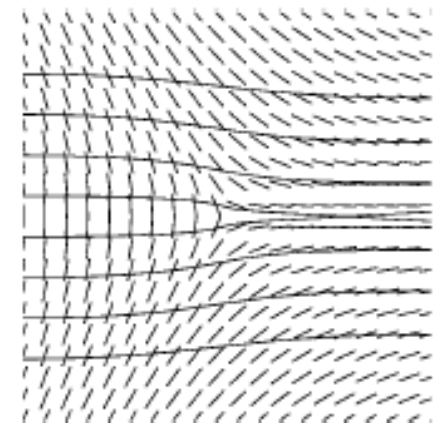
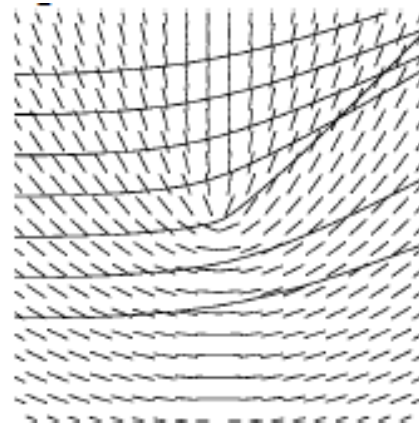
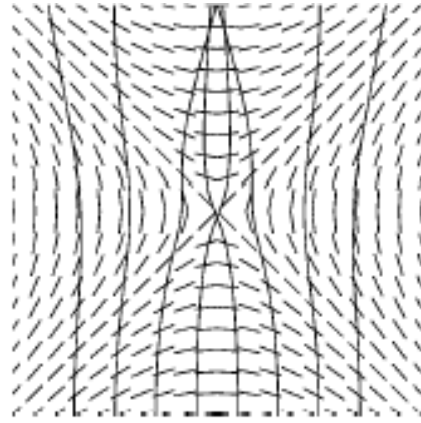
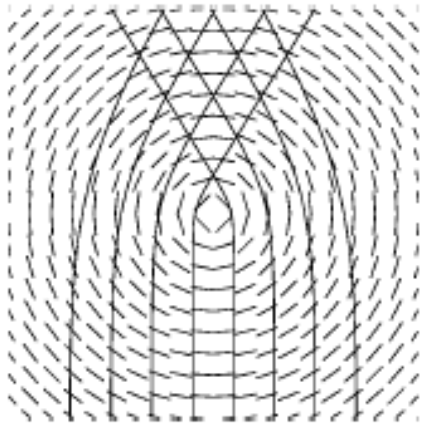
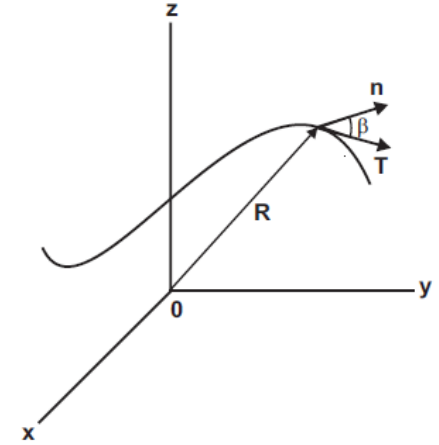


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C Satrio, F Moraes. EPJ E 20 (2006)

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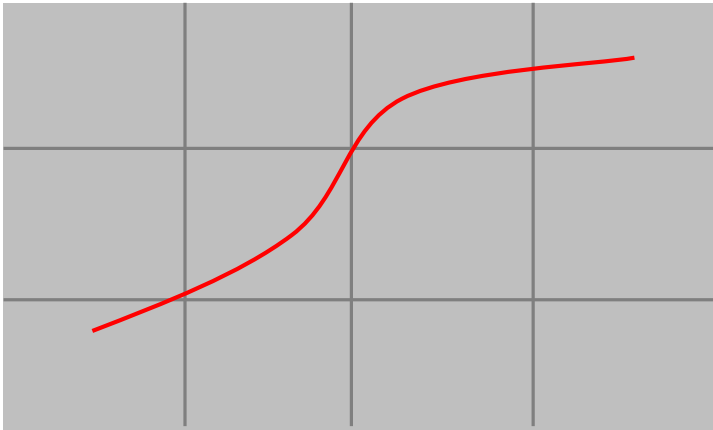
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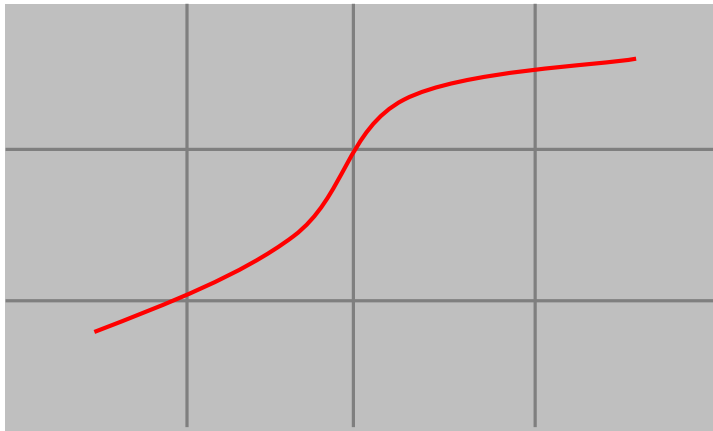


Curved path in a Euclidean space

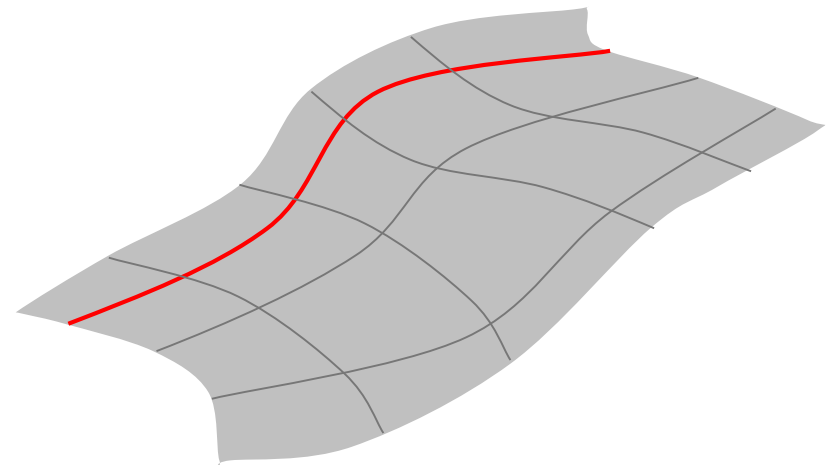
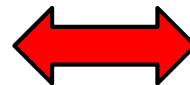
Fermat-Grandjean principle (1919)

- Extraordinary light paths obey a least action principle

$$\delta \left[\int_A^B N_e(\mathbf{r}) dl \right] = 0 = \delta \left[\int_A^B \sqrt{\varepsilon_{\perp} \cos^2 \beta(\mathbf{r}) + \varepsilon_{\parallel} \sin^2 \beta(\mathbf{r})} dl \right]$$



Curved path in a Euclidean space



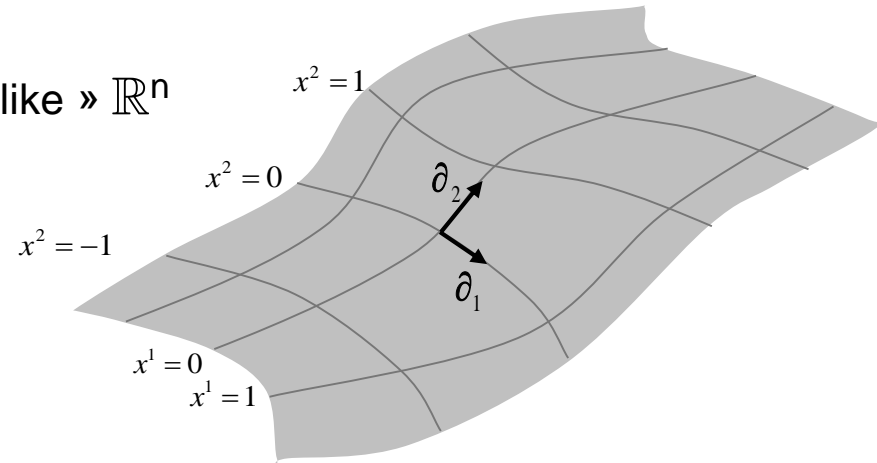
« Straight path » in a curved space

Reminder of Riemann geometry

► How to describe a curved geometry ?

- A **n-manifold** = smooth hypersurface that locally « looks like » \mathbb{R}^n

- coordinate system x^μ
- coordinate basis ∂_μ



- A **metric structure** to measure lengths

$$g_{\mu\nu} = \partial_\mu \cdot \partial_\nu$$

$$ds^2 = d\mathbf{x} \cdot d\mathbf{x} = (dx^\mu \partial_\mu) \cdot (dx^\nu \partial_\nu) \\ = dx^\mu dx^\nu g_{\mu\nu}$$

Generalized
 Pythagoras' theorem

- A **connection** to perform covariant differentiation... $\nabla_\alpha V^\mu = \partial_\alpha V^\mu + \Gamma_{\alpha\nu}^\mu V^\nu$

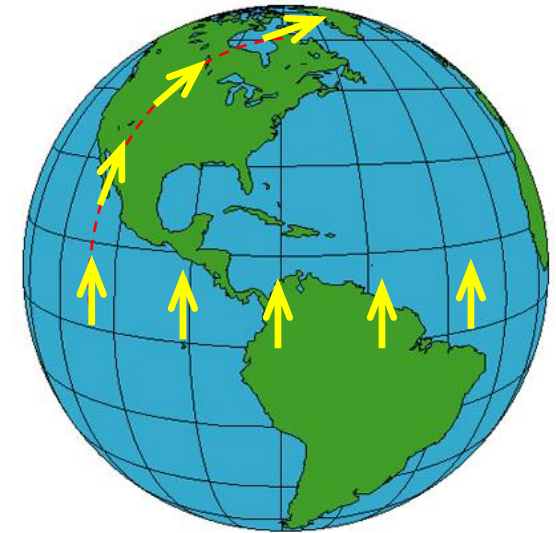
Levi-Civita connection

Reminder of Riemann geometry

... and to *parallel-transport* vectors, tensors, spinors ...

$$\begin{aligned}\frac{DT^{\mu..}_{v..}}{d\lambda} &= \frac{dx^\alpha}{d\lambda} \nabla_\alpha T^{\mu..}_{v..} = 0 \\ &= \frac{dT^{\mu..}_{v..}}{d\lambda} + \Gamma_{\alpha\beta}^\mu T^{\beta..}_{v..} \frac{dx^\alpha}{d\lambda} + \dots - \Gamma_{\alpha\nu}^\beta T^{\mu..}_{\beta..} \frac{dx^\alpha}{d\lambda} - \dots\end{aligned}$$

Searching for the North pole
with a compass



Reminder of Riemann geometry

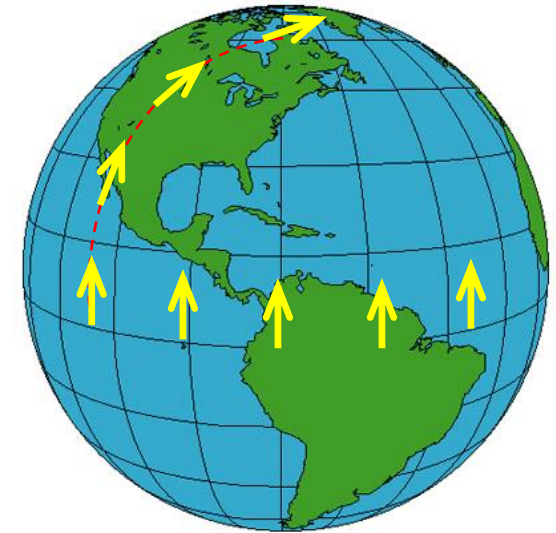
... and to *parallel-transport* vectors, tensors, spinors ...

$$\begin{aligned} \frac{DT^{\mu..}_{v..}}{d\lambda} &= \frac{dx^\alpha}{d\lambda} \nabla_\alpha T^{\mu..}_{v..} = 0 \\ &= \frac{dT^{\mu..}_{v..}}{d\lambda} + \Gamma^\mu_{\alpha\beta} T^{\beta..}_{v..} \frac{dx^\alpha}{d\lambda} + \dots - \Gamma^\beta_{\alpha\nu} T^{\mu..}_{\beta..} \frac{dx^\alpha}{d\lambda} - \dots \end{aligned}$$

- Parallel transport can be used to define a special class of curves, the **geodesics**, which are the curved-geometry generalizations of the Euclidean notion of straight lines. A geodesic curve is one that parallel-transport its own tangent vector (= autoparallel curve):

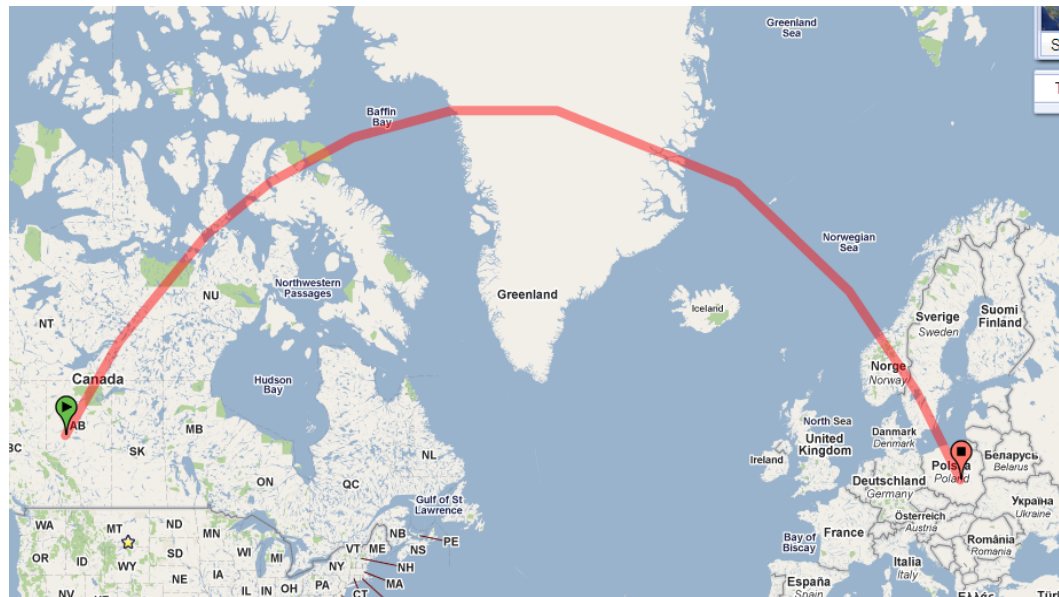
$$\frac{D}{d\lambda} \left(\frac{dx^\mu}{d\lambda} \right) = 0 = \frac{dx^\alpha}{d\lambda} \left(\frac{\partial}{\partial x^\alpha} \left[\frac{dx^\mu}{d\lambda} \right] + \Gamma^\mu_{\alpha\nu} \frac{dx^\nu}{d\lambda} \right) \Rightarrow 0 = \frac{d^2 x^\alpha}{d\lambda^2} + \Gamma^\mu_{\alpha\nu} \frac{dx^\nu}{d\lambda} \frac{dx^\alpha}{d\lambda} \quad \leftrightarrow \text{Newton's 2nd law}$$

Searching for the North pole
with a compass



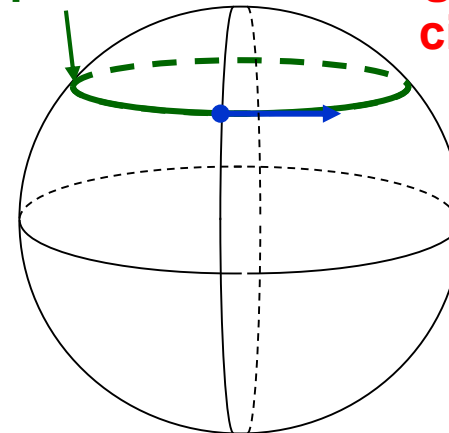
Reminder of Riemann geometry

- But most importantly, geodesics are the curves of extremal lengths (cf. Fermat-Grandjean principle). A warning: in the presence of curvature, actual geodesics may be very counter-intuitive:

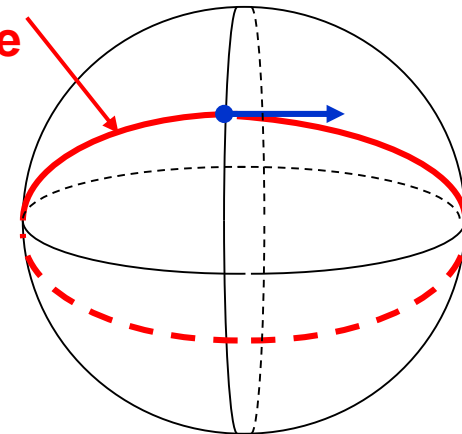


Shortest path between Calgary and Warsaw (flight plan)

parallel
great circle



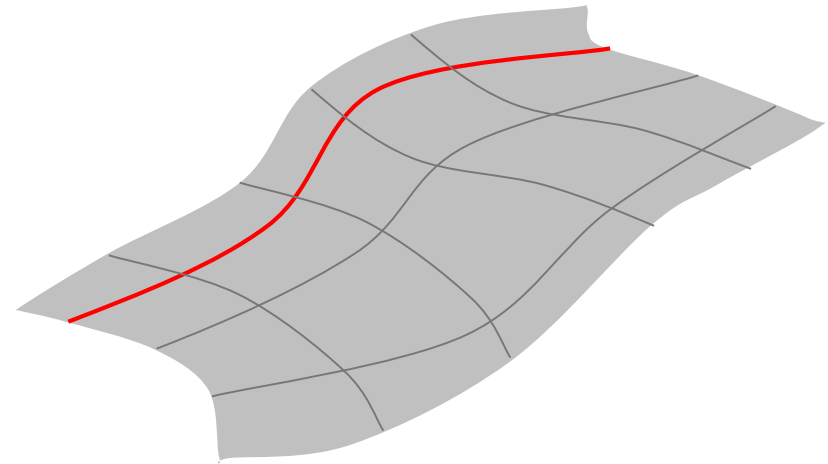
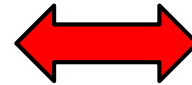
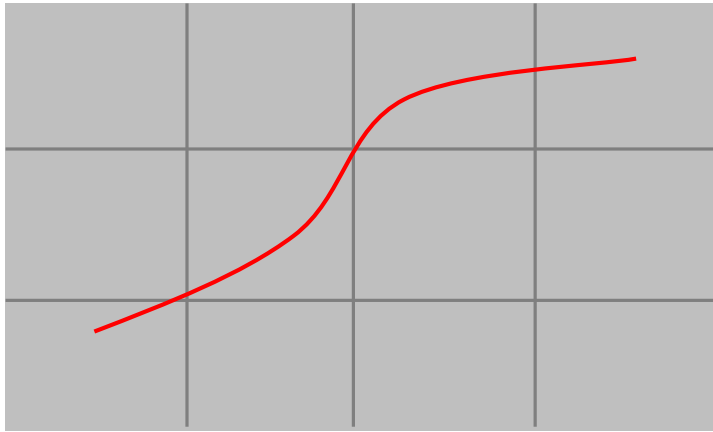
Not a geodesic



Geodesic

Transformation optics

U Leonhardt, TG. Philbin. Progress in Optics 53 (2009)
H Chen et al. Nature materials, 9 (2010)



Light path inside matter in Euclidean space

Geodesic in vacuum in curved space

$$ds_{2D}^2 = N_e^2(\mathbf{r}) dl^2$$

$$ds_{2D}^2 = g_{ij} dx^i dx^j$$

⇒ How to find the metric tensor representing a defective liquid crystal ?

Recipe for the line element

1. **Express the tangent vector T (planar path)**

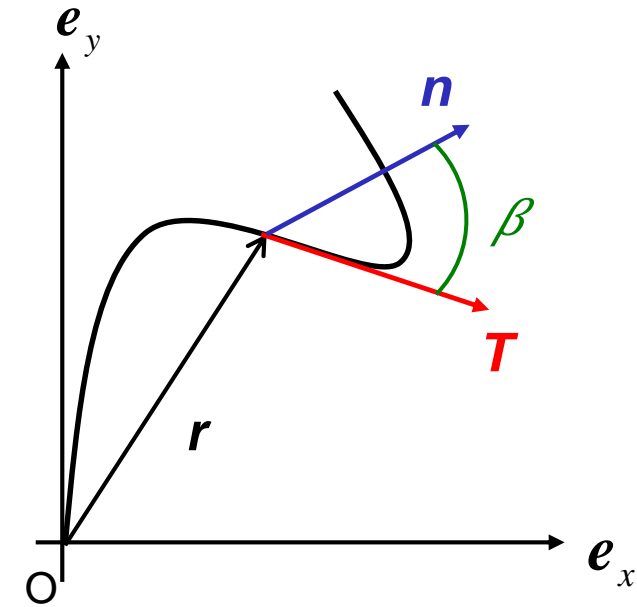
$$\mathbf{r} = r \cos \theta \mathbf{e}_x + r \sin \theta \mathbf{e}_y$$

$$\mathbf{T} = \frac{d\mathbf{r}}{dl} = (\dot{r} \cos \theta - r \dot{\theta} \sin \theta) \mathbf{e}_x + (\dot{r} \sin \theta + r \dot{\theta} \cos \theta) \mathbf{e}_y$$

2. **Write the components of the director field $\mathbf{n} = \cos \psi \mathbf{e}_x + \sin \psi \mathbf{e}_y$**

$$\Rightarrow \cos \beta = \mathbf{n} \cdot \mathbf{T} = \dot{r} \cos(\psi - \theta) + r \dot{\theta} \sin(\psi - \theta)$$

$$\sin \beta = \|\mathbf{n} \times \mathbf{T}\| = -\dot{r} \sin(\psi - \theta) + r \dot{\theta} \cos(\psi - \theta)$$



3. **Replace in the Fermat-Grandjean line element and see the magic**

$$N_e^2(\mathbf{r}) dl^2 = (\varepsilon_{\perp} \cos^2 \beta(\mathbf{r}) + \varepsilon_{\parallel} \sin^2 \beta(\mathbf{r})) dl^2$$

Recipe for the line element

- To spare repelling calculations, one sticks to $m=1$, $\psi_0=0$, but the proof is in the same fashion for the general case.

$$\Rightarrow \cos \beta = \dot{r} \cos(\psi - \theta) + r\dot{\theta} \sin(\psi - \theta) = \dot{r}$$

$$\sin \beta = -\dot{r} \sin(\psi - \theta) + r\dot{\theta} \cos(\psi - \theta) = r\dot{\theta}$$

$$\Rightarrow ds_{2D}^2 = N_e^2(\mathbf{r}) dl^2 = (\varepsilon_{\perp} \dot{r}^2 + \varepsilon_{\parallel} r^2 \dot{\theta}^2) dl^2 = \left(\varepsilon_{\perp} \left[\frac{dr}{dl} \right]^2 + \varepsilon_{\parallel} r^2 \left[\frac{d\theta}{dl} \right]^2 \right) dl^2 = \varepsilon_{\perp} dr^2 + \varepsilon_{\parallel} r^2 d\theta^2$$

A simple rescaling on the radial coordinate finally leads to

$$ds_{2D}^2 = d\rho^2 + \alpha^2 \rho^2 d\theta^2$$

Recipe for the line element

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$$\Rightarrow ds_{2D}^2 = N_e^2(\mathbf{r}) dl^2 = (\varepsilon_{\perp} \dot{r}^2 + \varepsilon_{\parallel} r^2 \dot{\theta}^2) dl^2 = \left(\varepsilon_{\perp} \left[\frac{dr}{dl} \right]^2 + \varepsilon_{\parallel} r^2 \left[\frac{d\theta}{dl} \right]^2 \right) dl^2 = \varepsilon_{\perp} dr^2 + \varepsilon_{\parallel} r^2 d\theta^2$$

A simple rescaling on the radial coordinate finally leads to

$$ds_{3D}^2 = d\rho^2 + \alpha^2 \rho^2 d\theta^2 + dz^2$$

⇒ What kind of geometry does this represent?

What a wedge cut does to space

- Ricci curvature scalar: $R(\rho) = \frac{(1-\alpha)}{\alpha\rho} \delta(\rho)$ = flat everywhere but on the z-axis.

For a circle of unit radius about the z-axis, the perimeter is given by $p = \oint_{\rho=1} ds_{3D} = \alpha \oint d\theta = 2\pi\alpha$

\Rightarrow « conical » geometry

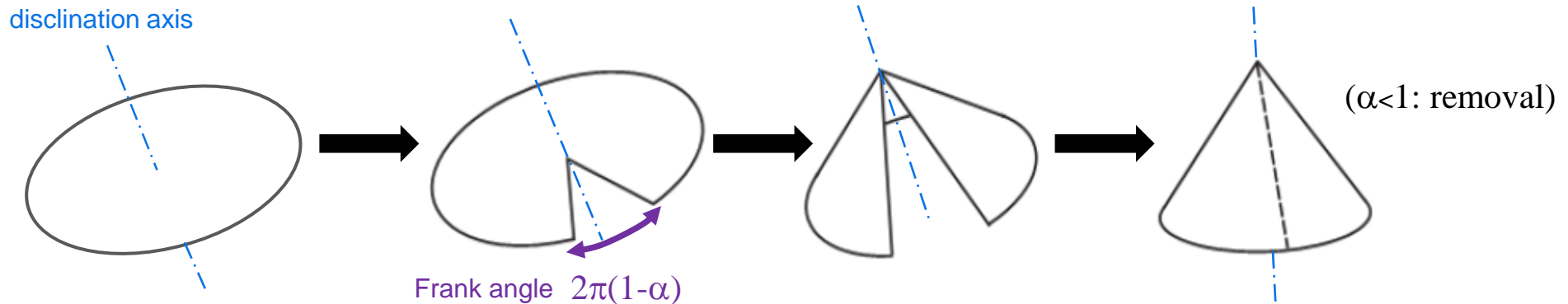
What a wedge cut does to space

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⇒ « conical » geometry

- Volterra process for a deficit-angle or wedge disclination:

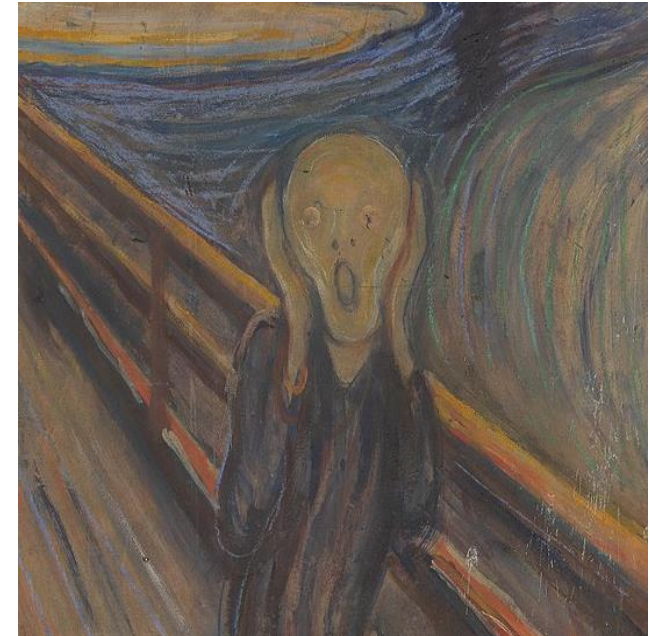


Just for fun...

- The general line element for a straight disclination of any topological charge is

$$ds_{3D}^2 = \left(\varepsilon_{\perp} \cos^2 \left[(m-1)\theta + \psi_0 \right] + \varepsilon_{\parallel} \sin^2 \left[(m-1)\theta + \psi_0 \right] \right) dr^2 \\ + \left(\varepsilon_{\perp} \sin^2 \left[(m-1)\theta + \psi_0 \right] + \varepsilon_{\parallel} \cos^2 \left[(m-1)\theta + \psi_0 \right] \right) r^2 d\theta^2 \\ - \left(\varepsilon_{\parallel} - \varepsilon_{\perp} \right) \sin \left[2(m-1)\theta + 2\psi_0 \right] r dr d\theta + dz^2$$

What should be done next involves computing the connection coefficients, the Riemann curvature tensor...



Distribution of defects

S Fumeron et al. Eur. Phys. J. B 90 (2017)

- In real-life, defects are not isolated. Sometimes, it is possible to find analytical expressions for the line element:

Discrete distribution of disclinations $ds^2 = -c^2 dt^2 + e^{-4V(x,y)} (dx^2 + dy^2) + dz^2$

$$V(x, y) = \frac{|F|}{4\pi} \ln \left[\left(\frac{\cosh^2 \left(\frac{\pi}{2a} (y - b) \right) - \cos^2 \left(\frac{\pi x}{2a} \right)}{\cosh^2 \left(\frac{\pi}{2a} (y - b) \right) - \sin^2 \left(\frac{\pi x}{2a} \right)} \right) \left(\frac{\cosh^2 \left(\frac{\pi}{2a} (y + b) \right) - \sin^2 \left(\frac{\pi x}{2a} \right)}{\cosh^2 \left(\frac{\pi}{2a} (y + b) \right) - \cos^2 \left(\frac{\pi x}{2a} \right)} \right) \right]$$

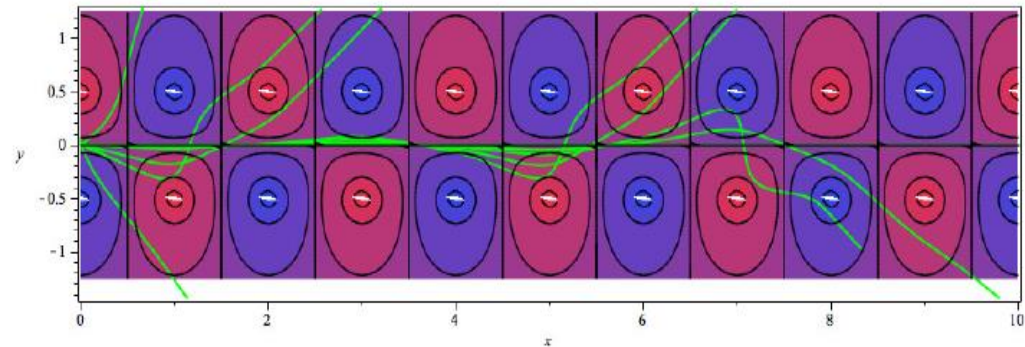
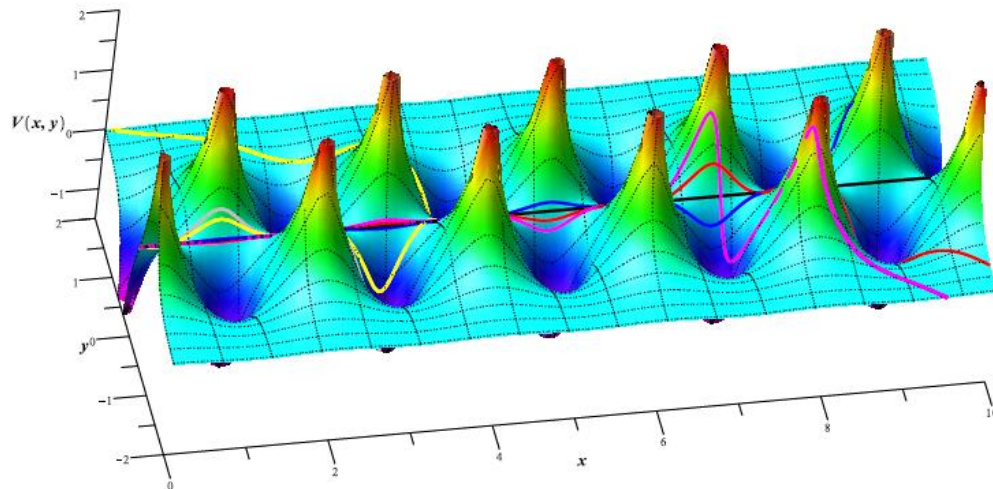
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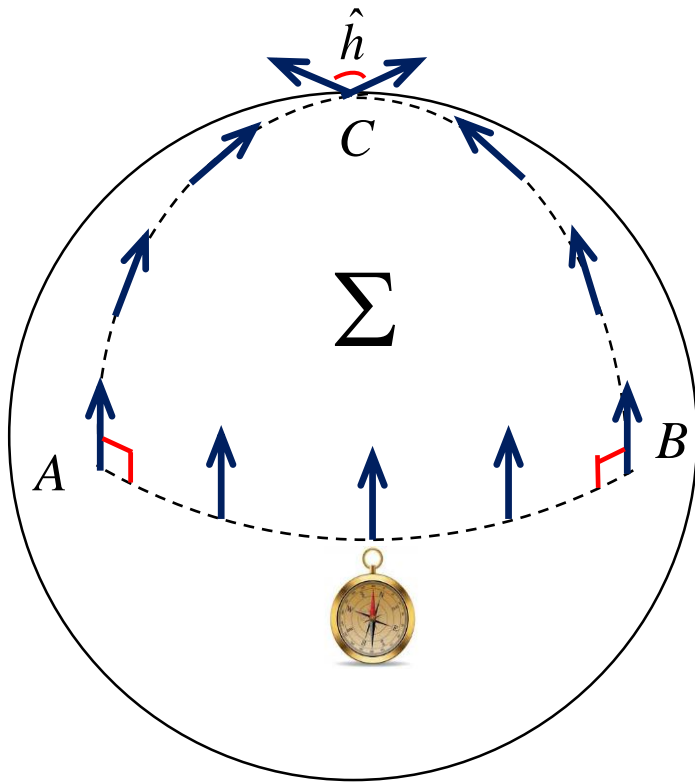
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Holonomy: intuitive approach

- Let us go back to « lost traveller problem »: after a closed loop, a parallel-transported vector fails to recover its initial direction: this is called (an) **holonomy**. How to understand that ?



Searching for the North pole

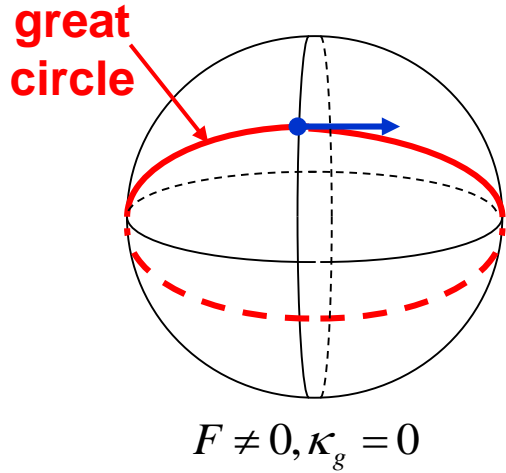
⇒ **Girard's formula** $\Sigma = R^2 \Omega = R^2 (\hat{A} + \hat{B} + \hat{C} - \pi)$

Here, this simplifies into $F\Sigma = \hat{C}$, which also turns out to be \hat{h} , the mismatch angle.

Hence, the mismatch angle is a measure of the Gaussian curvature F of the surface Σ bounded by the closed circuit.

This result is an basic outcome of the **Ambrose-Singer theorem**: for a given connection on a vector bundle, the curvature corresponds to the surface density of holonomy.

Holonomy: intuitive approach



$F = 0, \kappa_g \neq 0$

- But the Gaussian curvature F is also related to the topology of the surface:

Gauss-Bonnet theorem

$$\iint_{\Sigma} F dS + \oint_{\partial\Sigma} \kappa_g ds = 2\pi\chi$$

Gauss curvature (local)
 \leftrightarrow
 GEOMETRY

Geodesic curvature (local)
 \leftrightarrow
 GEOMETRY

Euler characteristic (global)
 \leftrightarrow
 TOPOLOGY

$$\Rightarrow \hat{h} = 2\pi\chi - \oint_{\partial\Sigma} \kappa_g ds$$

Therefore, holonomy is also connected to topology.

Holonomy: formal approach

S Carroll. Spacetime and geometry (2003)



H Hertz

- More generally, one defines **holonomy** as the failure to transport any information (such as the orientation of a vector) on a closed circuit on a curved surface Σ .
- Let there be a path parametrized by λ along which a vector V is parallel-transported. The *parallel propagator* Π is defined as

$$V^\mu(\lambda) = \Pi^\mu{}_\rho(\lambda) V^\rho(0)$$



E Cartan

- But the transport parallel condition also writes as: $\frac{DV^\nu}{d\lambda} = 0 \Rightarrow \frac{dV^\nu}{d\lambda} = -\Gamma^\nu_{\sigma\mu} \frac{dx^\sigma}{d\lambda} V^\mu = A^\nu{}_\mu V^\mu$

Therefore the parallel propagator obeys

$$\frac{d}{d\lambda} \Pi^\nu{}_\rho(\lambda) V^\rho(0) = A^\nu{}_\mu(\lambda) \Pi^\mu{}_\rho(\lambda) V^\rho(0) \quad \Rightarrow \quad \frac{d}{d\lambda} \Pi^\nu{}_\rho(\lambda) = A^\nu{}_\mu(\lambda) \Pi^\mu{}_\rho(\lambda)$$

Holonomy: formal approach

S Carroll. Spacetime and geometry (2003)

- This differential equation formally integrates as $\Pi^\mu_\rho(\lambda) = \delta^\mu_\rho + \int_0^\lambda A^\mu_\sigma(\eta) \Pi^\sigma_\rho(\eta) d\eta$

Similarly to what is done when establishing Dyson's formula (QFT), one iterates the process:

$$\begin{aligned} \Pi^\mu_\rho(\lambda) &= \delta^\mu_\rho + \int_0^\lambda A^\mu_\rho(\eta) d\eta + \int_{\eta_2=0}^\lambda \int_{\eta_1=0}^{\eta_2} A^\mu_\sigma(\eta_2) A^\sigma_\rho(\eta_1) d\eta_1 d\eta_2 + \dots \\ &+ \int_{\eta_3=0}^\lambda \int_{\eta_2=0}^{\eta_3} \int_{\eta_1=0}^{\eta_2} A^\mu_\sigma(\eta_3) A^\sigma_\nu(\eta_2) A^\nu_\rho(\eta_1) d\eta_1 d\eta_2 d\eta_3 + \dots \end{aligned}$$

- How to simplify this unpleasant formula?
 - ⇒ Instead of integrating over n -simplices, one integrates over n -cubes while keeping the product in the right order

$$\int_{\eta_n=0}^\lambda \int_{\eta_{n-1}=0}^{\eta_n} \dots \int_{\eta_1=0}^{\eta_2} A(\eta_n) A(\eta_{n-1}) \dots A(\eta_1) d^n \eta = \frac{1}{n!} \int_{\eta_n=0}^\lambda \int_{\eta_{n-1}=0}^{\eta_n} \dots \int_{\eta_1=0}^{\eta_2} P[A(\eta_n) A(\eta_{n-1}) \dots A(\eta_1)] d^n \eta$$

Holonomy: formal approach

S Carroll. Spacetime and geometry (2003)

- Thanks to Taylor's expansion series, the previous formula « miraculously » simplifies into

$$\Pi(\lambda) = I + \sum_{n=1}^{+\infty} \frac{1}{n!} \int_{\eta_n=0}^{\lambda} \int_{\eta_{n-1}=0}^{\lambda} \dots \int_{\eta_1=0}^{\lambda} P[A(\eta_n)A(\eta_{n-1})\dots A(\eta_1)] d^n \eta = P \exp \int_0^{\lambda} A(\eta) d\eta$$

with P is the ordering operator. On a loop γ about a point M , the holonomy writes explicitly as

$$\Pi^{\mu}_{\nu}[\gamma] = P \exp \left(-\oint_{\gamma(M)} \Gamma^{\mu}_{\sigma\nu} \frac{dx^{\sigma}}{d\eta} d\eta \right) \Leftrightarrow \Pi[\gamma] = P \exp \left(-\oint_{\gamma(M)} \Gamma_{\sigma} dx^{\sigma} \right)$$

- Ambrose-Singer theorem** = to know the holonomy at every point of the manifold is equivalent to know the curvature at every point of the manifold \leftrightarrow quantum loop gravity.

Holonomy due to a disclination

AM de Carvalho, C Satiro, F Moraes. EPL 80 (2007)

- For a loop about the origin in a $z=C^{st}$ plane, only the polar connection symbol is retained

$$\Pi[\gamma] = P \exp\left(-\oint_{\gamma(M)} \Gamma_{\theta} d\theta\right)$$

$$\Gamma_{\theta} = \frac{m}{\alpha} \left(\alpha^2 \cos^2 [(m-1)\theta + \psi_0] + \sin^2 [(m-1)\theta + \psi_0] \right) \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix}$$

- For $ds_{2D}^2 = d\rho^2 + \alpha^2 \rho^2 d\theta^2$, one gets $\Gamma_{\theta} = \alpha \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix}$ and a Taylor expansion of the parallel propagator gives:

$$\Pi[\gamma] = \begin{bmatrix} \cos(2\pi\alpha) & -\sin(2\pi\alpha) \\ \sin(2\pi\alpha) & \cos(2\pi\alpha) \end{bmatrix}$$

Disclination \Leftrightarrow **Active rotation** = when acting on a vector, causes its counterclockwise rotation.

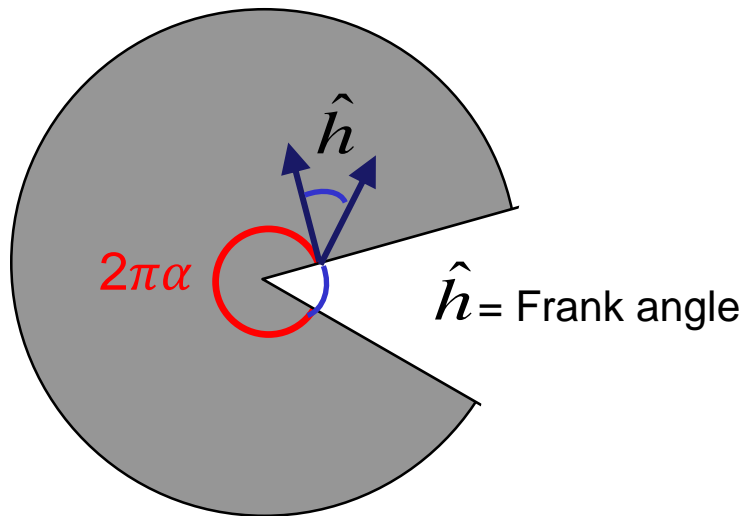
Holonomy due to a disclination

AM de Carvalho, C Satiro, F Moraes. EPL 80 (2007)

- How much does a vector turn when describing a loop around the defect ?

$$\hat{h} = 2\pi - 2\pi\alpha$$

Mismatch angle of a vector after a loop Loop term $\theta \in [0, 2\pi]$ Active rotation (counterclockwise)



- Gauss-Bonnet version (much more efficient than using Π):

$$\hat{h} = 2\pi\chi - \oint_{\partial\Sigma} \kappa_g ds = 2\pi - \int_{\theta=0}^{2\pi} \frac{1}{R} R\alpha d\theta = 2\pi(1 - \alpha)$$

$$ds^2 = \alpha^2 R^2 d\theta^2 \quad \text{circle of radius } R$$

Let us do the experiment !!

Comparison with Aharonov-Bohm

« *The Aharonov-Bohm effect is real physics not ideal physics* ».

Aharonov-Bohm phase

Magnetic flux is confined within the solenoid, it vanishes everywhere else.

It has measurable effects: shift of the electronic interference pattern...

Quantum

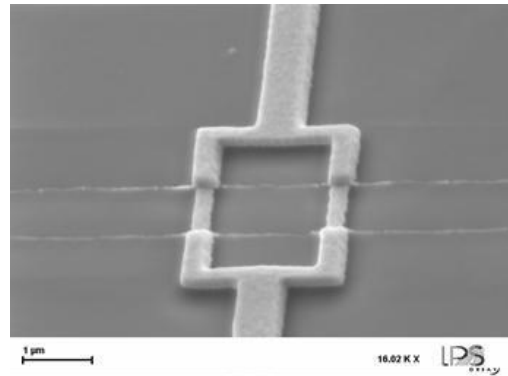
Moraes phase

Curvature is confined within the disclination line, it vanishes everywhere else.

It has measurable effects: rotation of polarization plane of linearly polarized light...

Classical

SQUID



Still waiting for its NDC device...

M Berry

A tentative definition

Phase	First appeared in	Mostly known in	Parameter space	Topological	Adiabatic
Pancharatnam	1956	Optics	Poincaré sphere	No	Yes
Aharonov-Bohm	1959	Quantum electrodynamics	Spacetime	Yes	No
Exchange statistics (of Abelian anyons)	1977 1982 1984	Condensed matter	Real space	Yes	Yes
Berry	1983 1984	Quantum mechanics	General	No	Yes
Aharonov-Casher	1984	Quantum electrodynamics	Real space	Yes	No
Hannay angle	1985	Classical mechanics	Real space	No	Yes
Aharonov-Anandan	1987	Quantum mechanics	General	Yes	No
Zak	1989	Condensed matter	Momentum space	No	No

They are all examples of what is generically called *geometric* or *Berry phases*, that is « **phases are not attributed to the forces applied onto the [quantum] system. Instead, they are associated with the connection of space itself.** »

A crash-course on liquids crystals
Topological defects in nematics
Topological defects everywhere?

Zoology of line defects
Geometry of line defects
Geometric phase

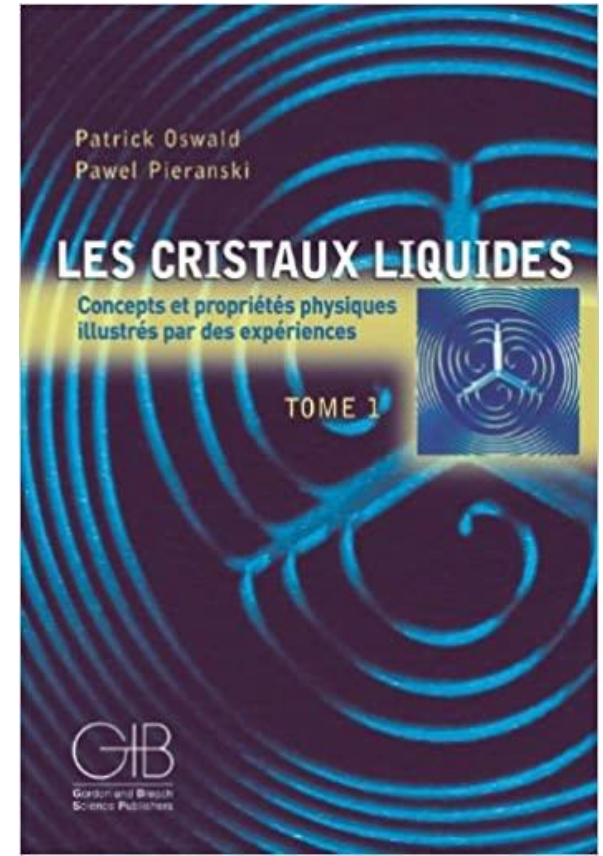
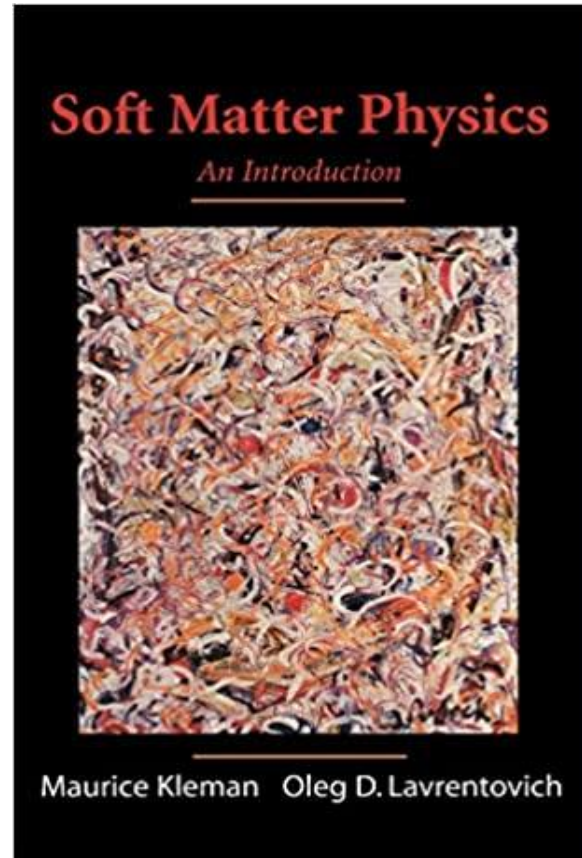
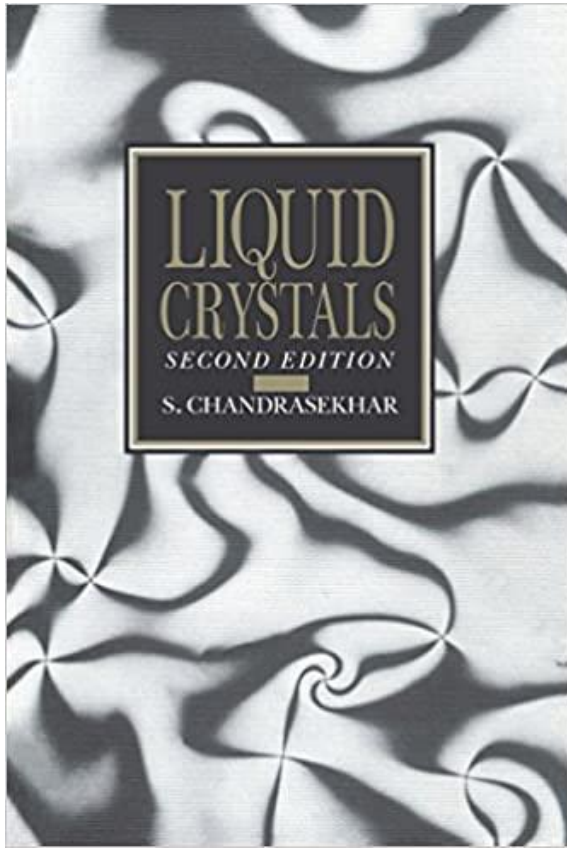
Next lecture...



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Appendices

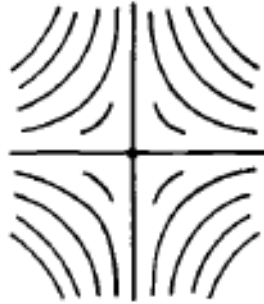
Disclination gallery



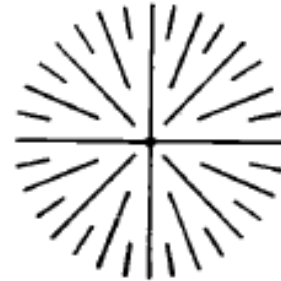
$s = \frac{1}{2}$



$s = -\frac{1}{2}$



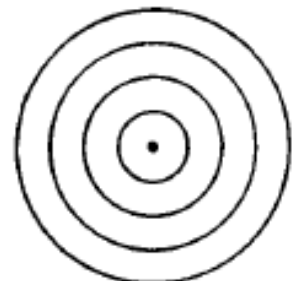
$s = -1$



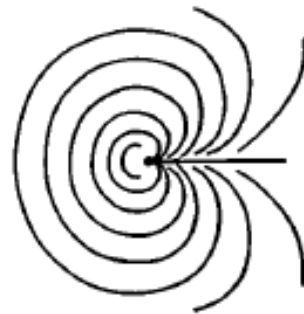
$s = 1, c = 0$



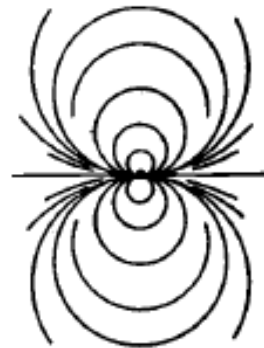
$s = 1, c = \pi/4$



$s = 1, c = \pi/2$

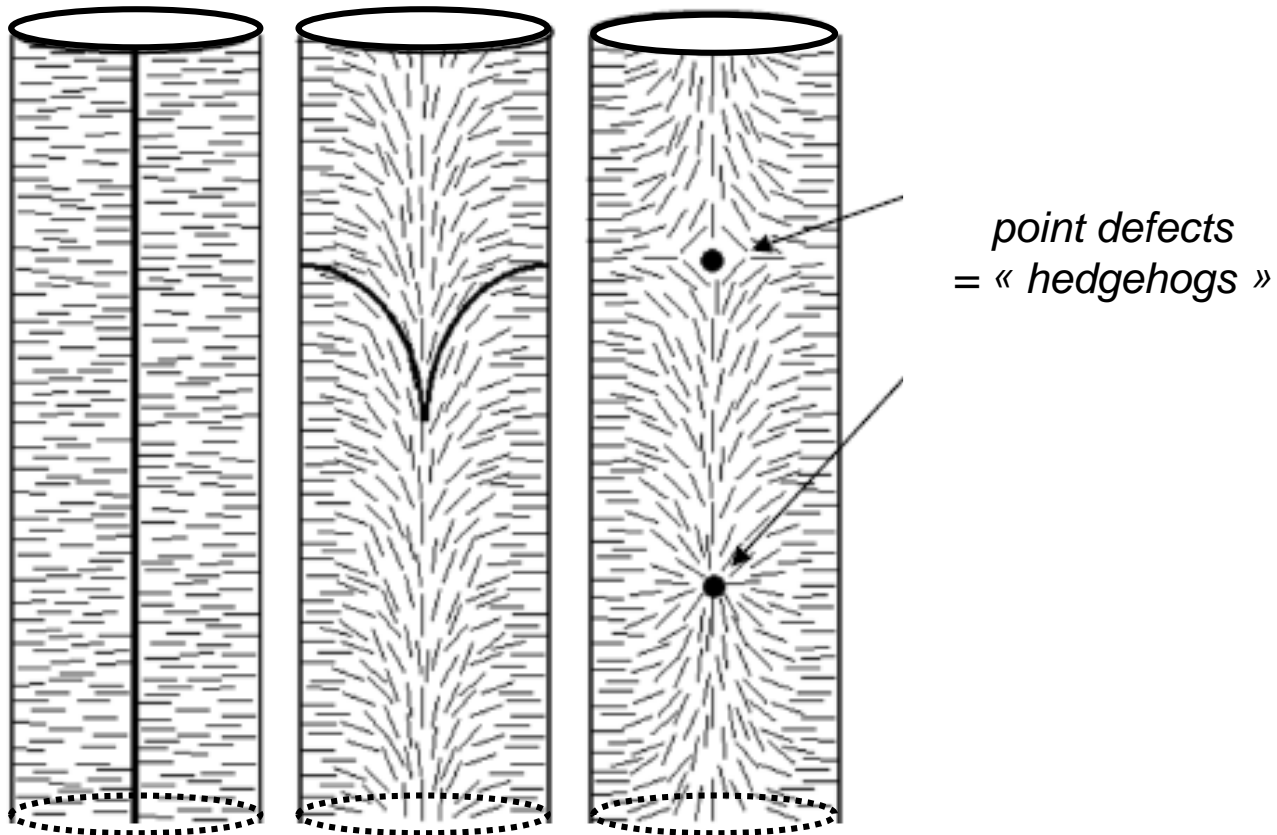


$s = 3/2$

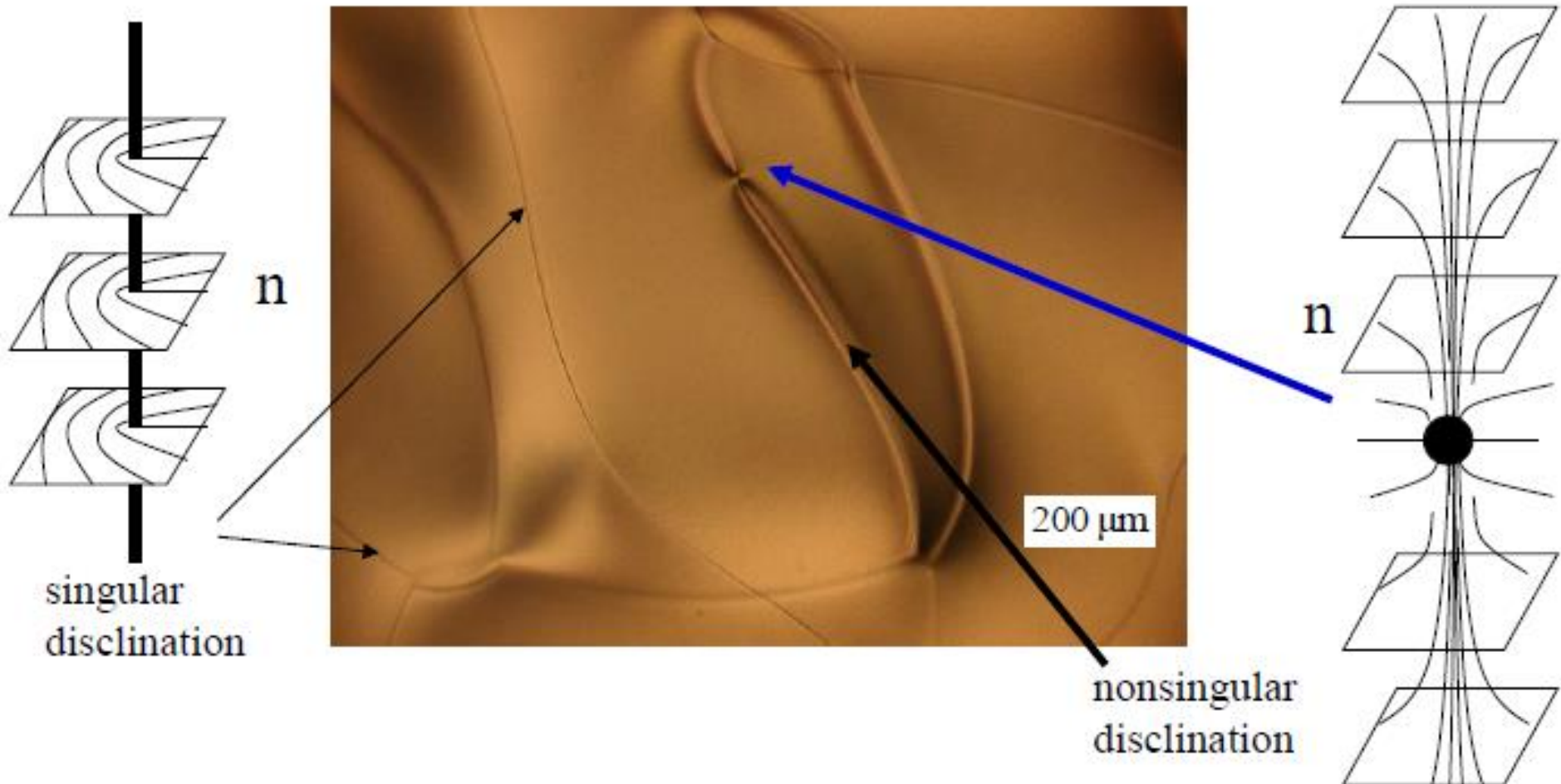


$s = 2$

Escape in the third dimension



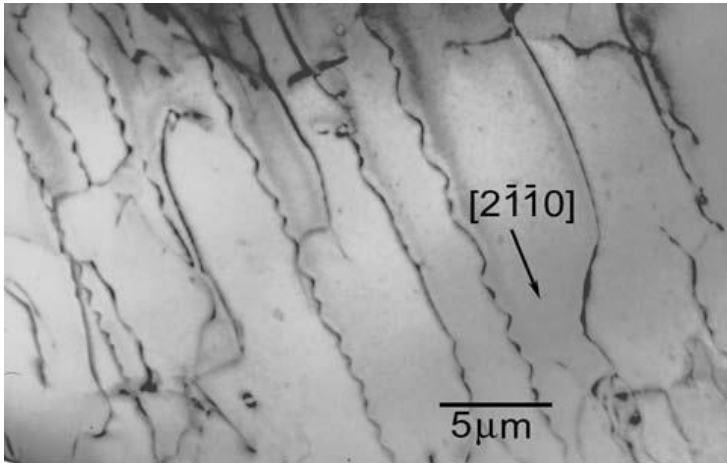
Escape in the third dimension



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Twisting things around

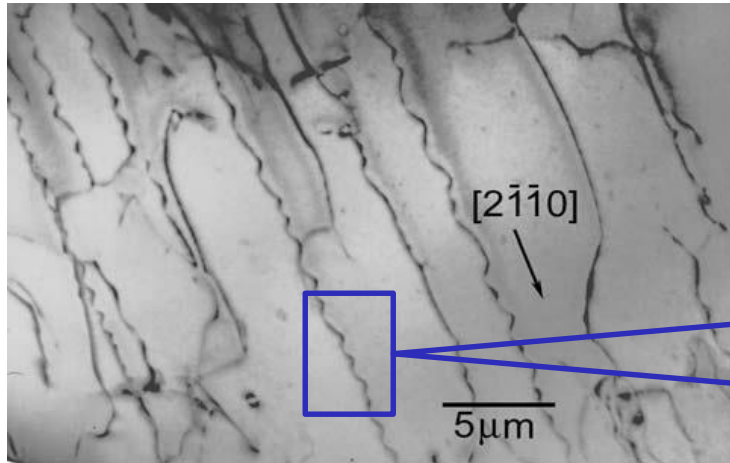
- Line defects can also have a chirality. For instance in crystalline solids, a very common chiral defect is the screw dislocation:



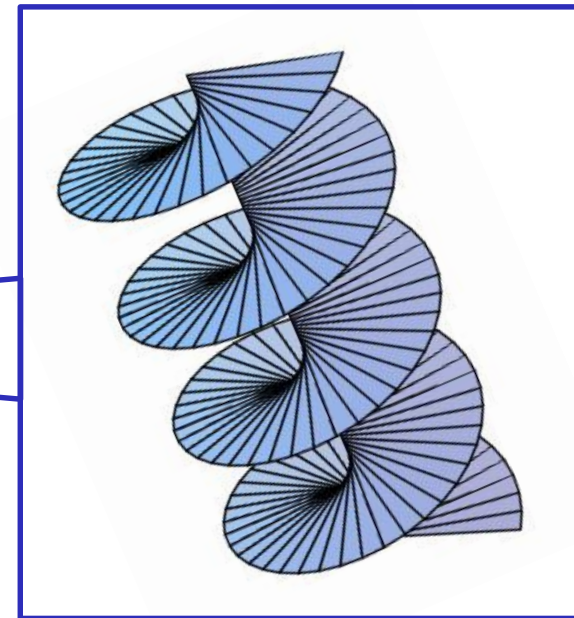
TEM image in CaCO_3

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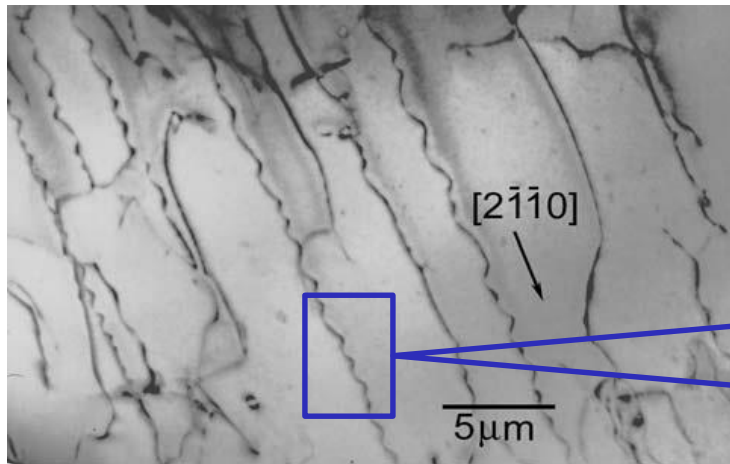


TEM image in CaCO_3



Twisting things around

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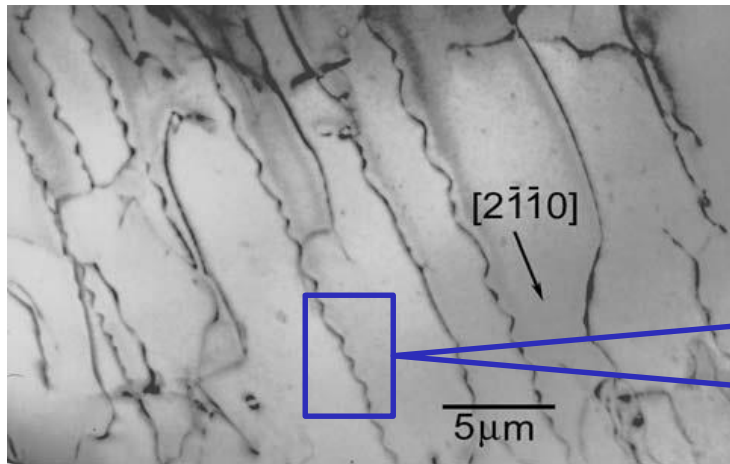
TEM image in CaCO_3



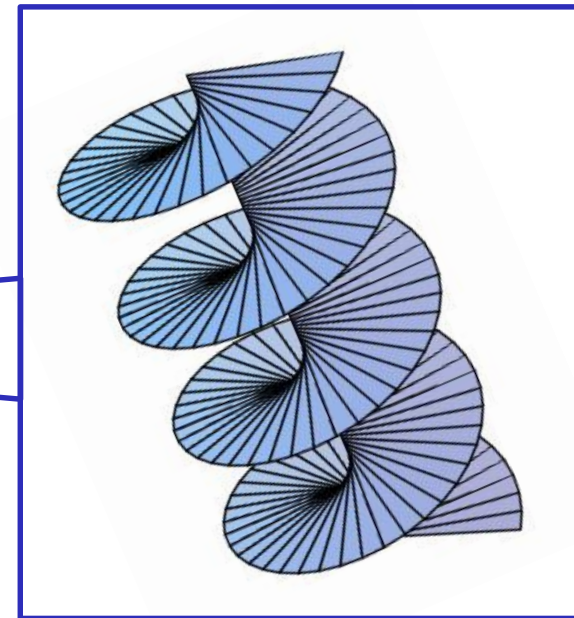
A « giant screw dislocation » in Da Vinci's spiral stairway (Chambord)

Twisting things around

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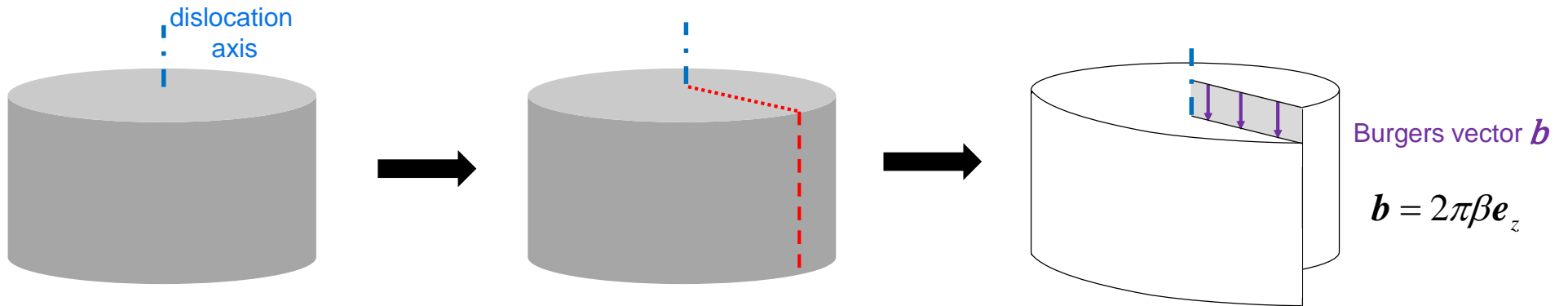
TEM image in CaCO_3



- The defect couples the rotation and translation around the z -axis: $\theta \rightarrow \theta + 2\pi \Leftrightarrow z \rightarrow z + 2\pi\beta$
The sign of β dictates if the helix is left-handed or right-handed.

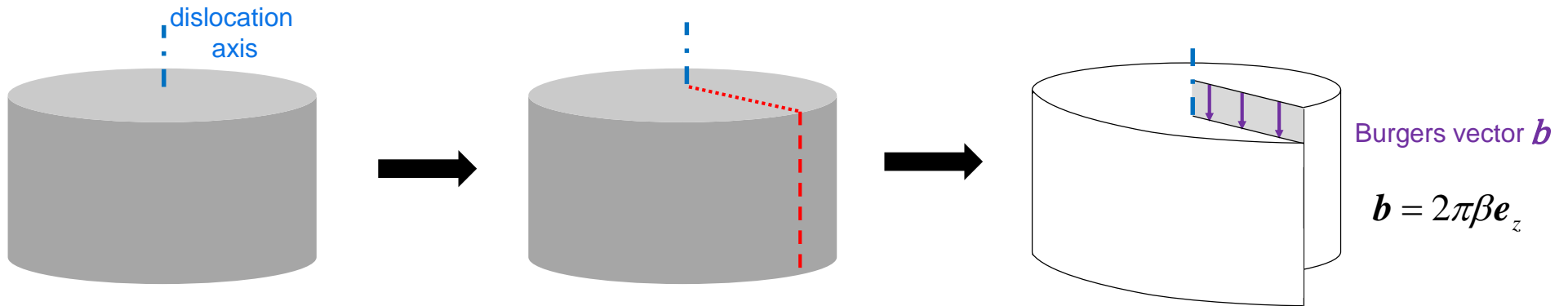
Twisting things around

- Volterra process for the screw dislocation:

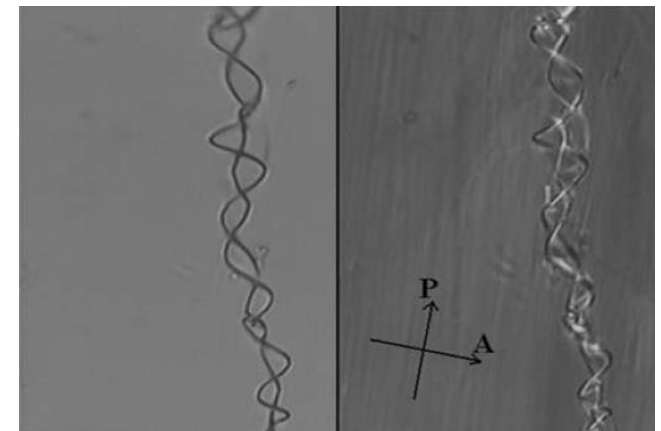


Twisting things around

- Volterra process for the screw dislocation:



- In mesophases as well, chiral line defects have been observed:



POM image in a smectic (C. Blanc, Montpellier 2)

Twisting things around

- The geometry of a single twisted disclination (or dispiration) is given by the line element:

$$ds_{3D}^2 = dr^2 + \alpha^2 r^2 d\theta^2 + (\beta d\theta + dz)^2$$

- Differential geometry of a dispiration is much harder than it may look at first sight:

1. *Distributional non-vanishing curvature on the z-axis:*

$$R(r) = \frac{(1-\alpha)}{\alpha r} \delta(r)$$

$V_4 =$ Riemann geometry

2. *Distributional non-vanishing torsion on the z-axis:*

$$T^z = 2\pi\beta \frac{\delta(r)}{r} dr \wedge d\theta$$

$U_4 =$ Riemann-Cartan geometry

Distribution of defects

- In real-life, defects are not isolated. Sometimes, it is possible to find analytical expressions for the line element:

1. *Discrete distribution of disclinations:* $ds^2 = -c^2 dt^2 + e^{-4V(x,y)} (dx^2 + dy^2) + dz^2$

$$V(x, y) = \frac{|F|}{4\pi} \ln \left[\left(\frac{\cosh^2 \left(\frac{\pi}{2a} (y - b) \right) - \cos^2 \left(\frac{\pi x}{2a} \right)}{\cosh^2 \left(\frac{\pi}{2a} (y - b) \right) - \sin^2 \left(\frac{\pi x}{2a} \right)} \right) \left(\frac{\cosh^2 \left(\frac{\pi}{2a} (y + b) \right) - \sin^2 \left(\frac{\pi x}{2a} \right)}{\cosh^2 \left(\frac{\pi}{2a} (y + b) \right) - \cos^2 \left(\frac{\pi x}{2a} \right)} \right) \right]$$

2. *Continuous distribution of dislocations:* $ds_{3D}^2 = dr^2 + r^2 d\theta^2 + (\Omega r^2 d\theta + dz)^2$

$$\Omega = Nb/2 \quad \text{Surface density of Burgers vector}$$

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$\Omega = Nb/2$ *Surface density of Burgers vector*

3. *Continuous distribution of chiral disclinations ($r > R$):*

$$ds_{3D}^2 = dr^2 + r^{2(A-1)} d\theta^2 + (\Omega r^2 d\theta + dz)^2 \quad A=1+qR^2/2 \quad \text{Surface density of deficit angle}$$